

Global oscillation modes in the evolving disk of B emission stars

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B emission (Be) stars : Show Balmer lines in emission.
Emission lines \rightarrow cool circumstellar disk

V/R variations: line profile variability caused by global disk oscillation ($m=1$) mode with period about several years.

More remarkable variability in disk formation stage than in disk dissipation stage.

See poster for more details

Evolving disk :

