

The Venus Climate Database VCD version 2.3 (and 3.0)

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and *the VCD Team*



A diagram showing a purple circle containing the text "VCD routine suite" in red and black. Below this, the function call `call_vcd(z_key,z,lon,lat,...,T,p,dens,zonwind,...)` is shown in black text.

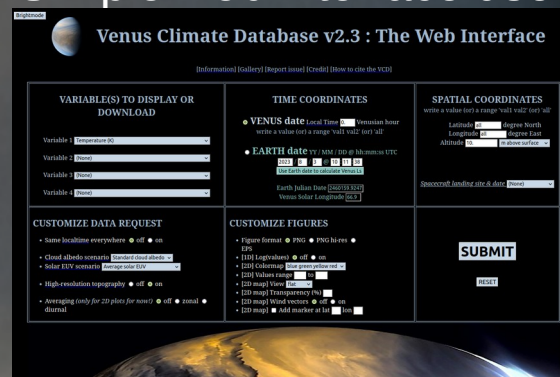
VCD
routine suite

`call_vcd(z_key,z,lon,lat,...,
T,p,dens,zonwind,...)`

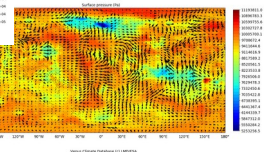
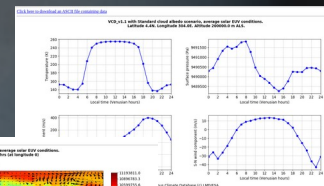
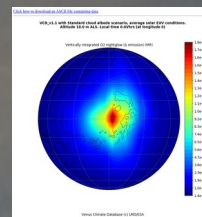
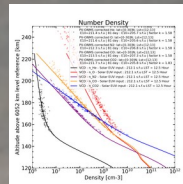
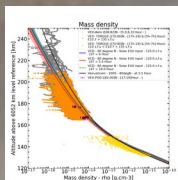
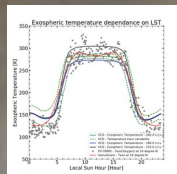
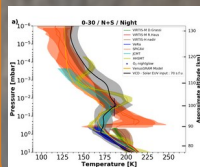
Direct use in users' codes

Various language interfaces

Simple web interface use

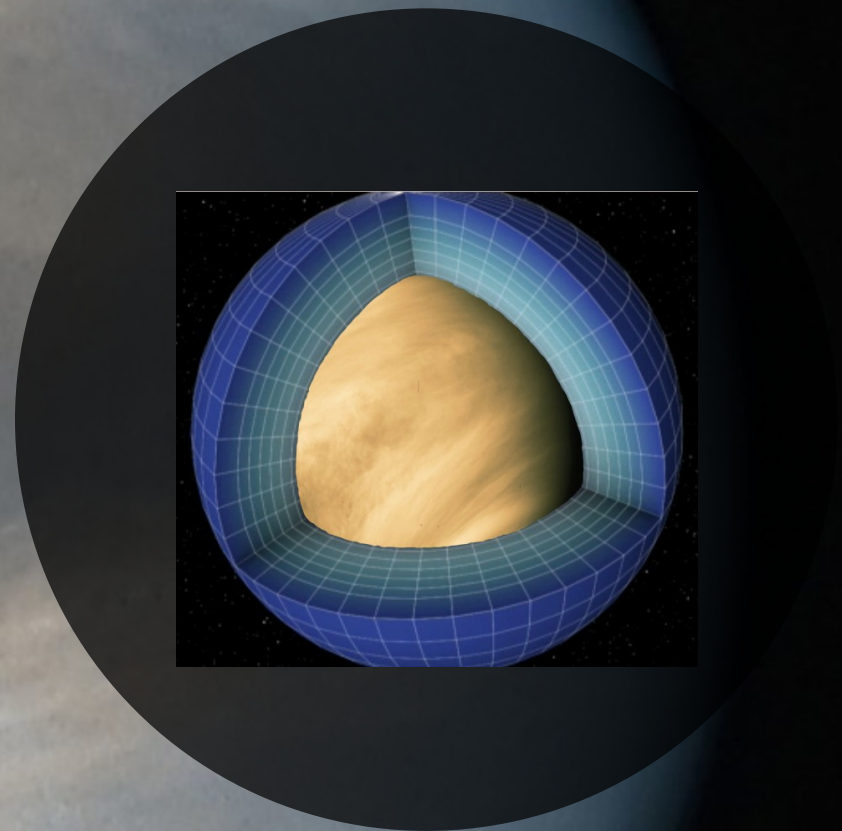


Validation with observational datasets



The Venus PCM

- Three-dimensional:
96x96x [50 (0~95 km) / 78 (0~150 km) / 90 (0~250 km)]
- Vertical coordinates: hybrid (sigma/pressure)
- Dynamical core, transport of tracers
- Specific physics:
 - ◆ Radiative transfer: Infrared Net Exchange Rates matrix
Solar heating rates: tables
 - ◆ Thermosphere: Non-LTE processes
EUV heating
molecular diffusion
 - ◆ Parameterizations of sub-grid processes:
boundary layer (Mellor&Yamada 1982), convection
non-orographic gravity waves
orographic gravity waves
 - ◆ Topography
- Includes photochemistry



Gilli et al (2017, 2021) ; Garate & Lebonnois (2018) ;
Navarro et al (2018, 2021) ;
Martinez et al (2023, 2024) ; Stolzenbach et al (2023)

The Venus Climate Database

- ESA is funding our Venus Climate Database in the context of the EnVision mission
- The VCD is intended to be useful for engineering applications (aerobraking ; Entry, Descent & Landing) and scientific work which requires accurate knowledge of the Venusian atmosphere (e.g. analysis of observations)
- The VCD is freely available, either via light online access for moderate use (web interface) or full version including advanced post-processing software

<https://www-venus.lmd.jussieu.fr>

VCD 2.2 was released in April 2023

VCD 2.3 was released in September 2023

Software

Scenarios

High
resolution
mode

Variability
and
perturbations

Examples of
validation

Examples of
applications

Various
interfaces

VCD Software

The main access software (call_vcd) :

Primarily composed of a Fortran routine designed to provide access, with adequate interpolations, to fields and variables as a result of a **point-wise (in location and time) query**.

Enabling the user to query along the time dimension either by specifying an **Earth date** or a **Venus Local Time**.

The VCD dataset includes a **full climatological Venus day**, built using 10 days of GCM simulations, sampled at 1/24th of a Venusian day to accurately represent the diurnal cycle.

VCD Scenarios

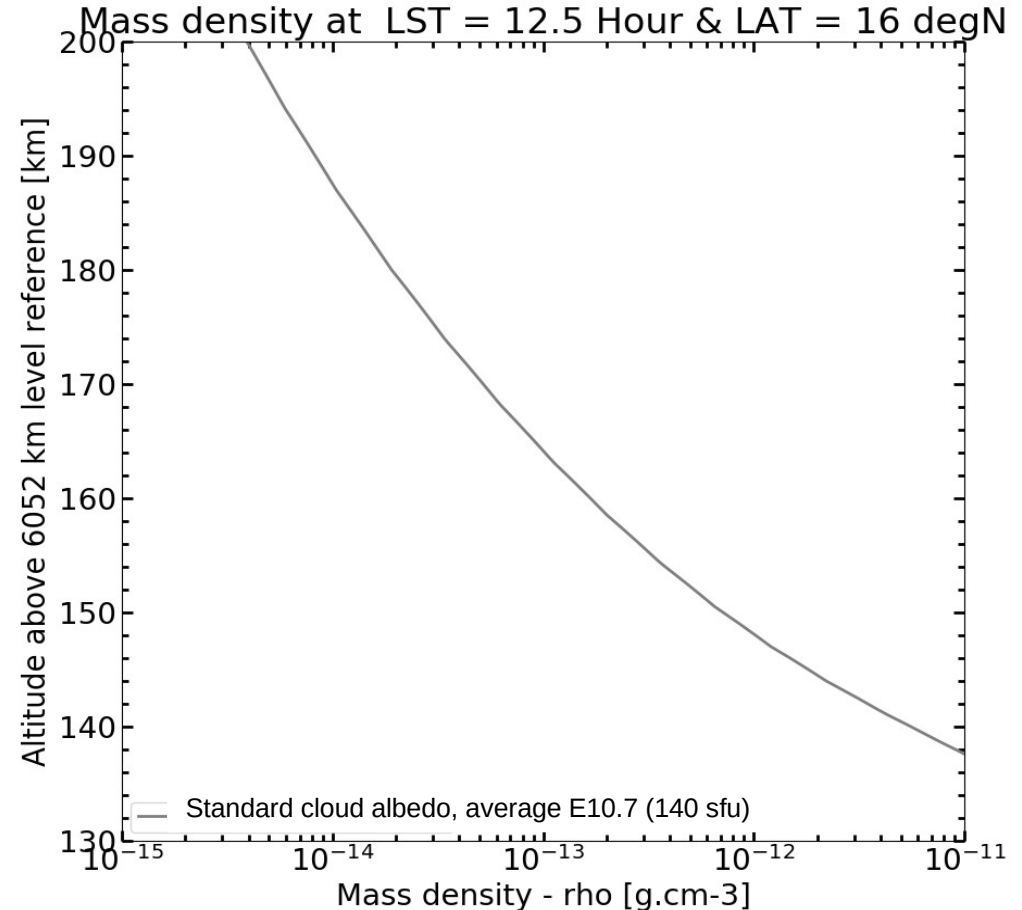
Scenarios

- **Cloud UV albedo** (standard, min or max)
- **EUV E10.7 index** (average, min, max, Earth date or user chosen value)

Example

VCD vs density measurements

- The **EUV scenarios** of the VCD allow to bracket reality
- The comparison can be more accurate using a **specific EUV** (interpolation)



VCD Scenarios

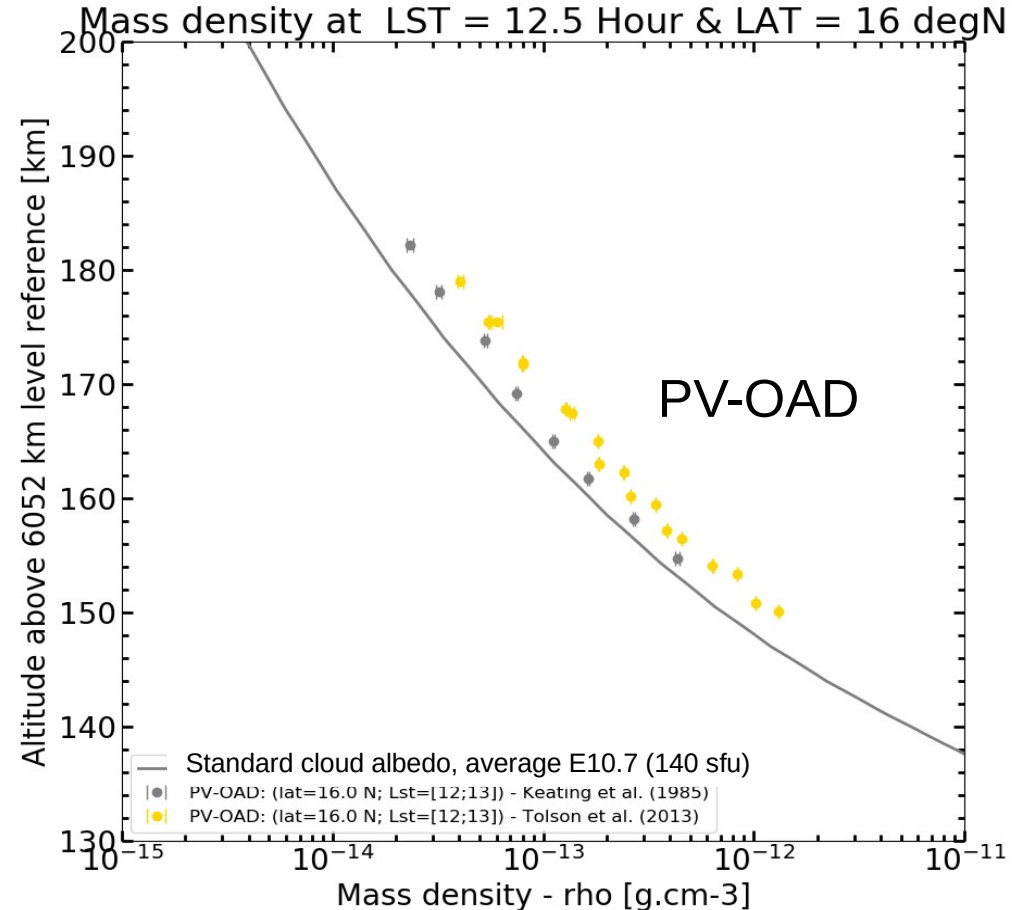
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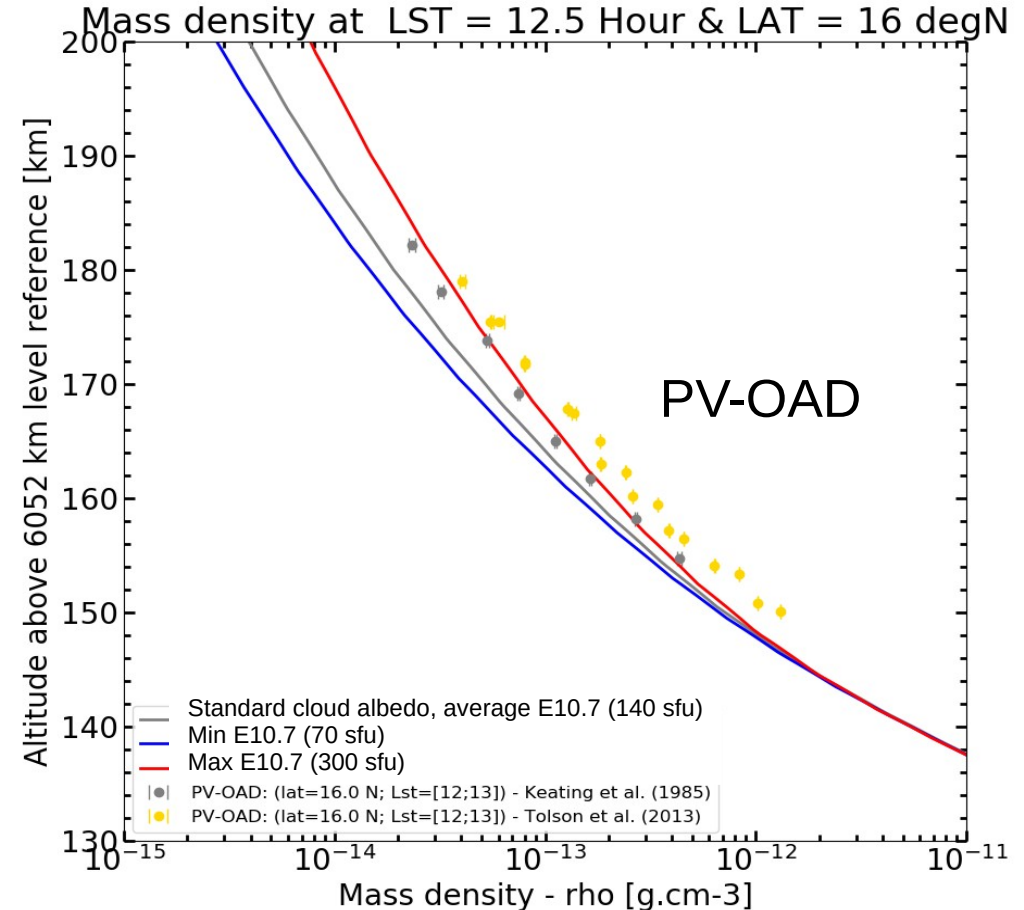
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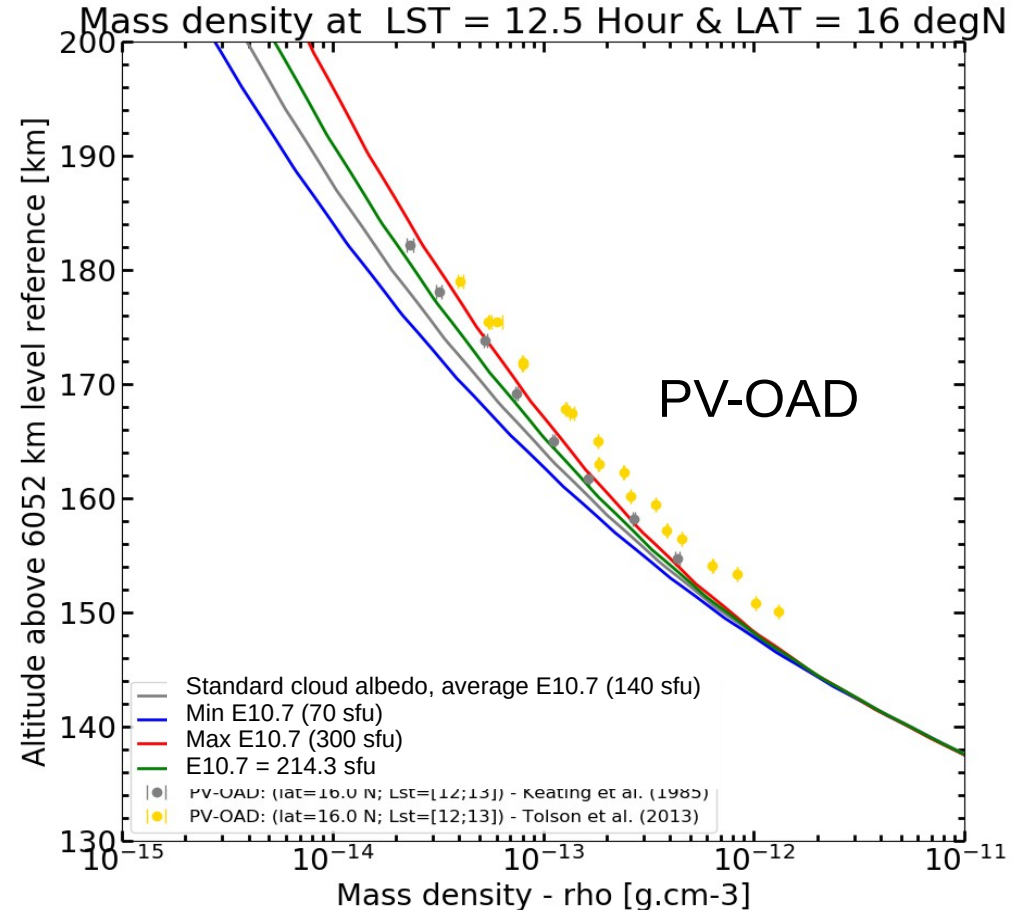
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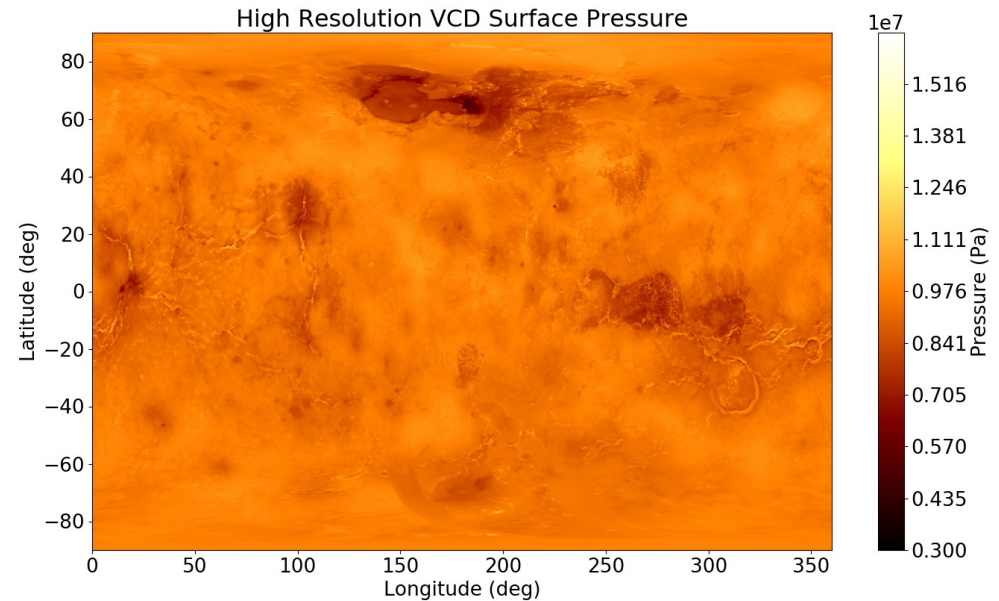
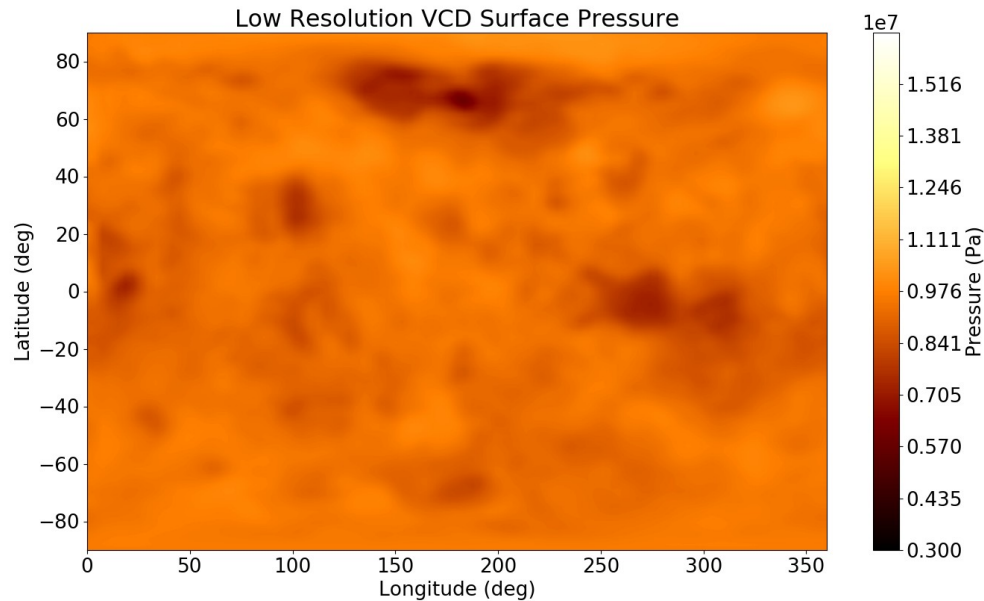
VCD vs density measurements

- The **EUV scenarios** of the VCD allow to bracket reality
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VCD High resolution mode

The VCD software include a “**high resolution**” mode using Magellan (**23 pixels/degree** map) topography combined to re-derived surface pressure and near-surface temperatures (scheme based on hydrostatic equilibrium and known lapse rate)

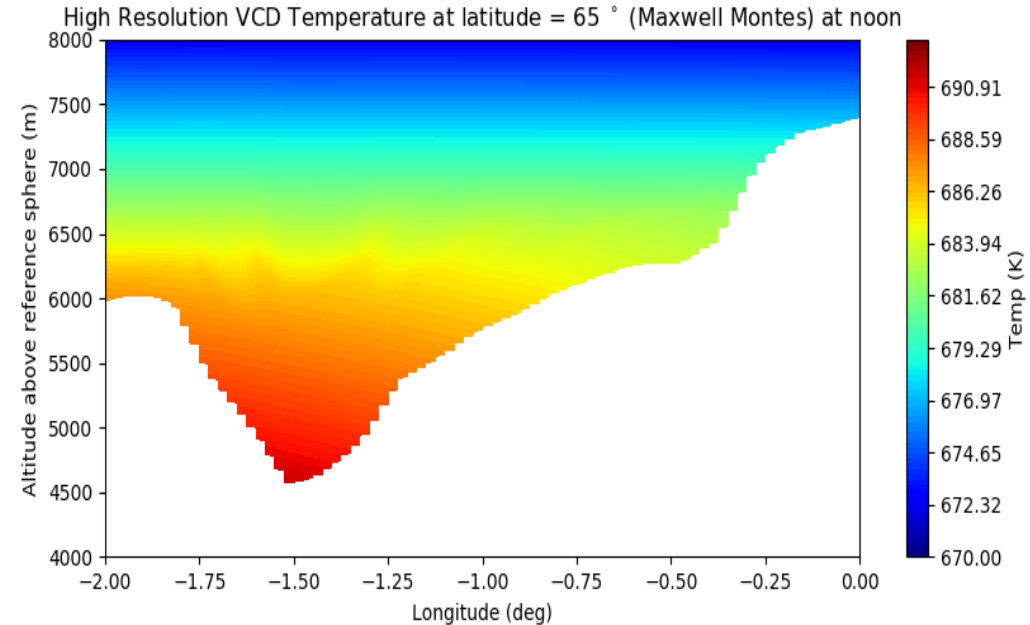
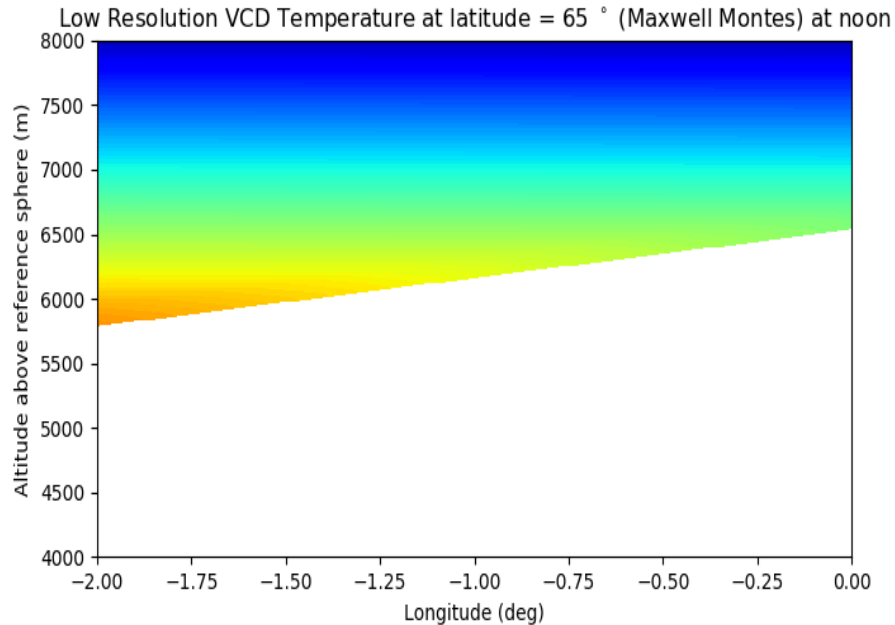


GCM resolution (PCM 96x96 grid) : Longitude x latitude : $3.75^\circ \times 1.875^\circ$

High resolution (Magellan map & PV measurements to fill gaps) : Longitude x latitude : $0.04^\circ \times 0.04^\circ$

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VCD Variability and perturbations schemes

Day-to-day
and V-hourly
RMS

Gravity
waves

Reconstruction
of the
meteorology
using EOFs

The VCD provides three types of variability

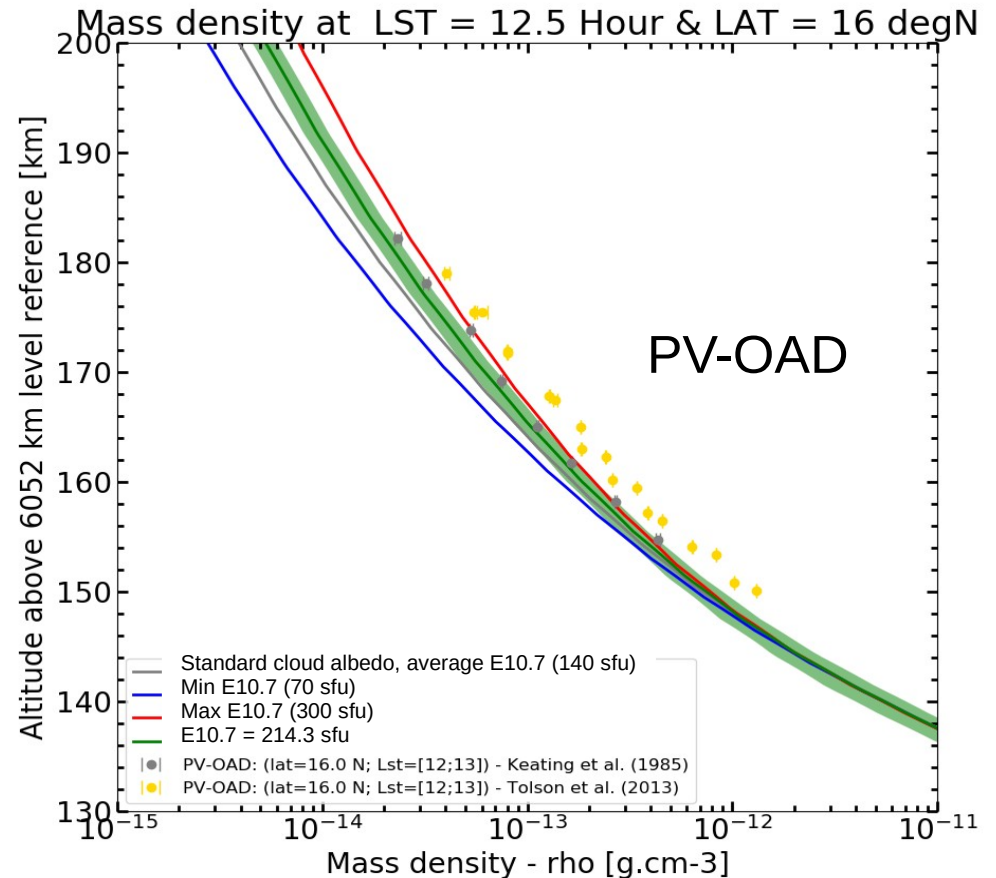
- **RMS variability from GCM simulations** for mean atmospheric fields (V-hourly and day-to-day)
- Possibility to add random **small-scale perturbations** as gravity waves (of user-specified wavelength)
- Possibility to add random **large-scale perturbations** (extracted from EOFs of individual GCM runs)

Variability: V-hourly and day-to-day RMS

Example

VCD vs density measurements

V-hourly RMS is shown as green area around the (E10.7=214.3 sfu) climatological mean density

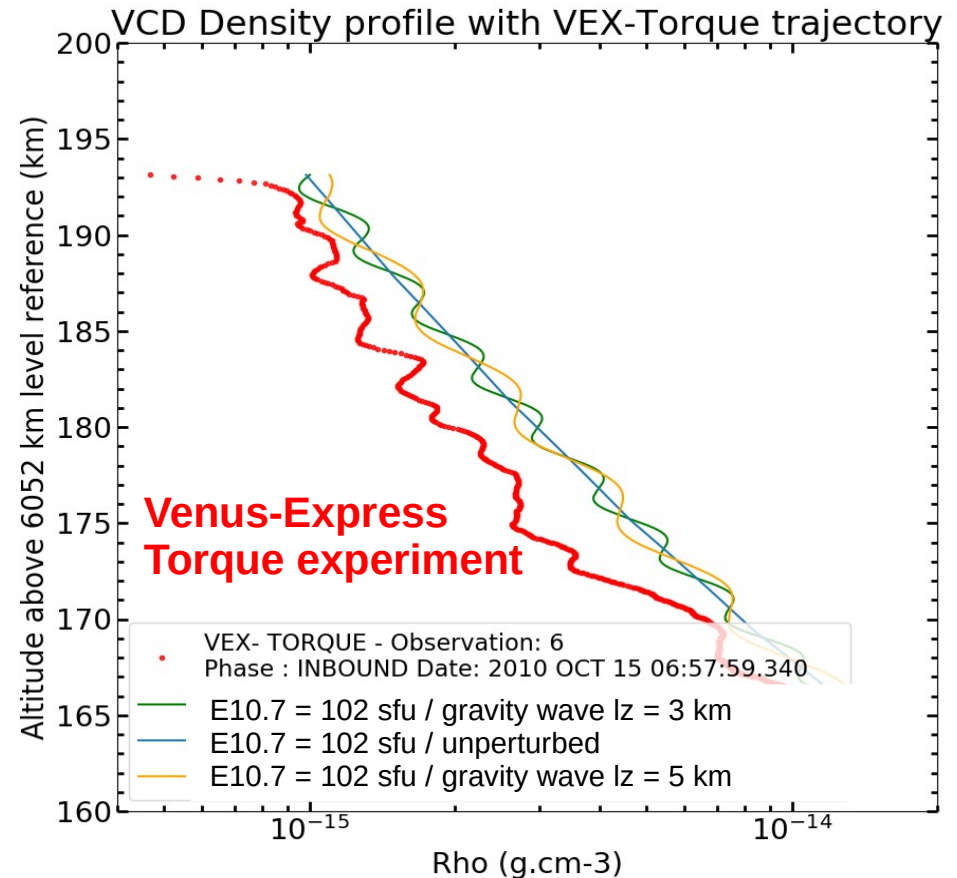


Perturbations: Small-scale gravity waves

Example

In addition to the climatology, the VCD allow to add **small-scale perturbations** (gravity waves generated at the top of the cloud convective layer).

The vertical wavelength is specified by the user. The horizontal wavelength is set to be $\lambda_h = 100 \lambda_z$

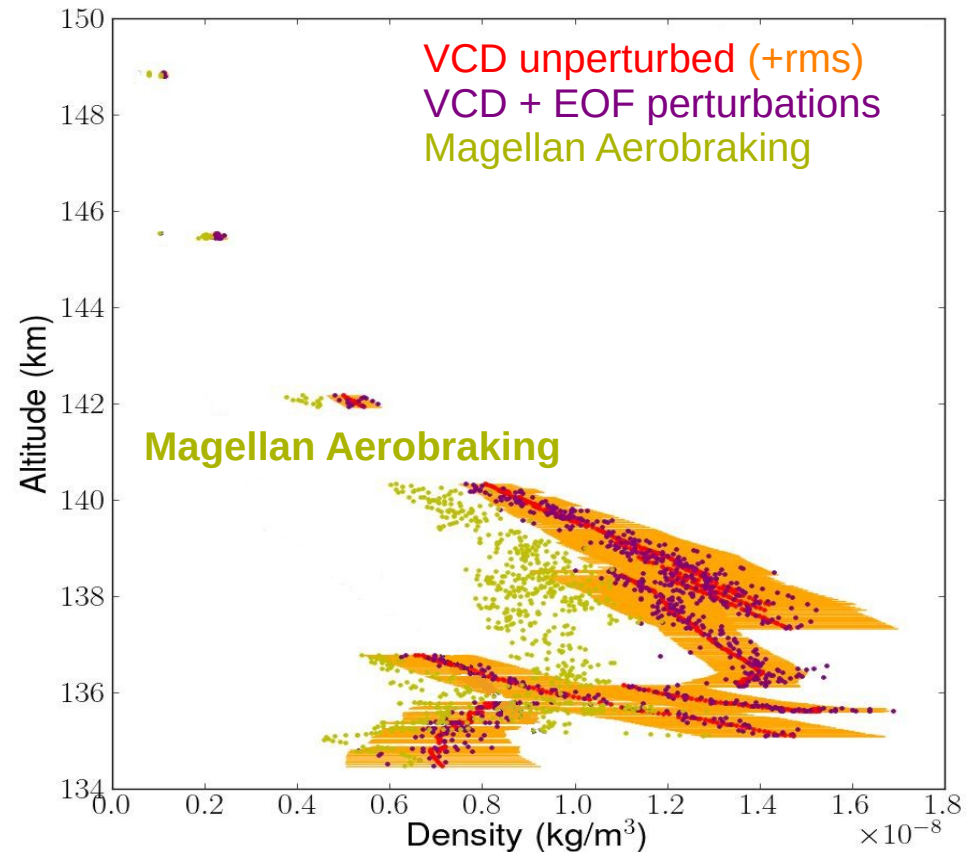


Perturbations: Large-scale, using EOF reconstruction

Example

In addition to the climatology (red dots), the VCD allow to add **large-scale perturbations** to reconstruct meteorology, using **EOFs of high-temporal resolution GCM simulations** (purple dots).

(orange area is previously mentioned RMS)



Examples of VCD validation

VCD Validation based on datasets from

- Pioneer Venus (OAD, ONMS),
- Magellan (aerobraking, POD)
- Venus-Express (VeXADE),

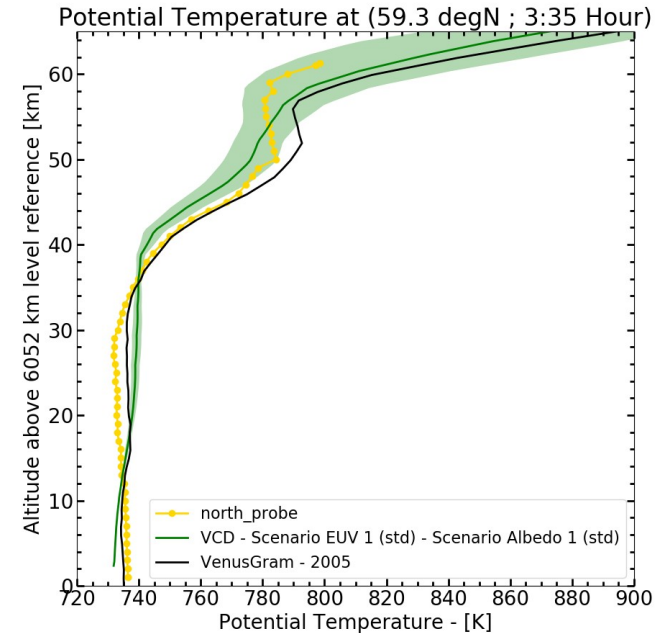
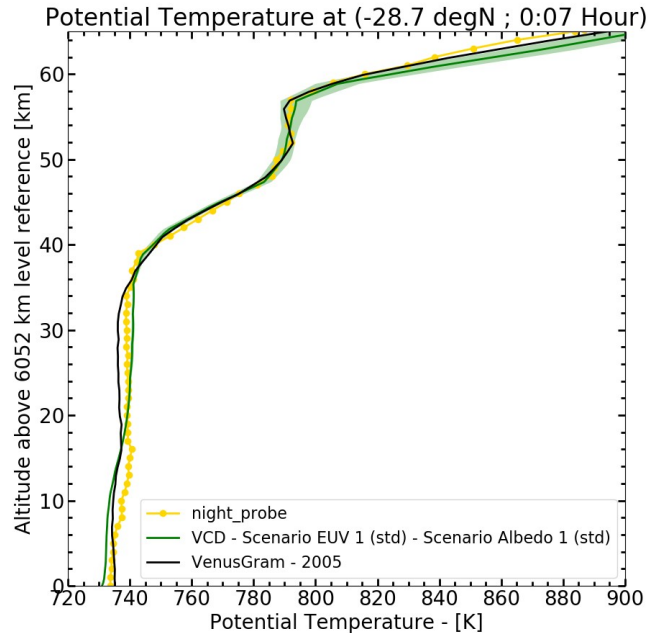
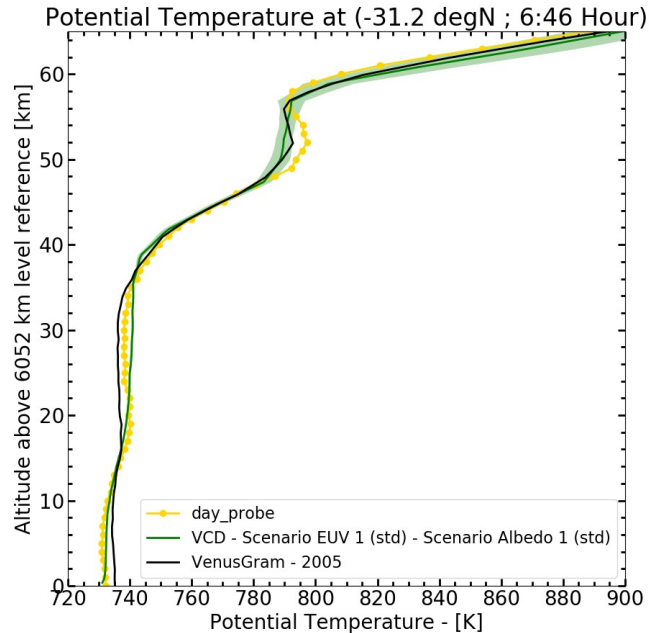
we can investigate temperature, density and composition above 150 km.
=> check the **VCD Validation Document**

Note that the PCM tuning includes :

- EUV efficiency and CO_2 -O quenching coefficient
- non-orographic gravity waves
- O number density => major scientific question

Parameter to be taken into account for thermospheric cases : E10.7

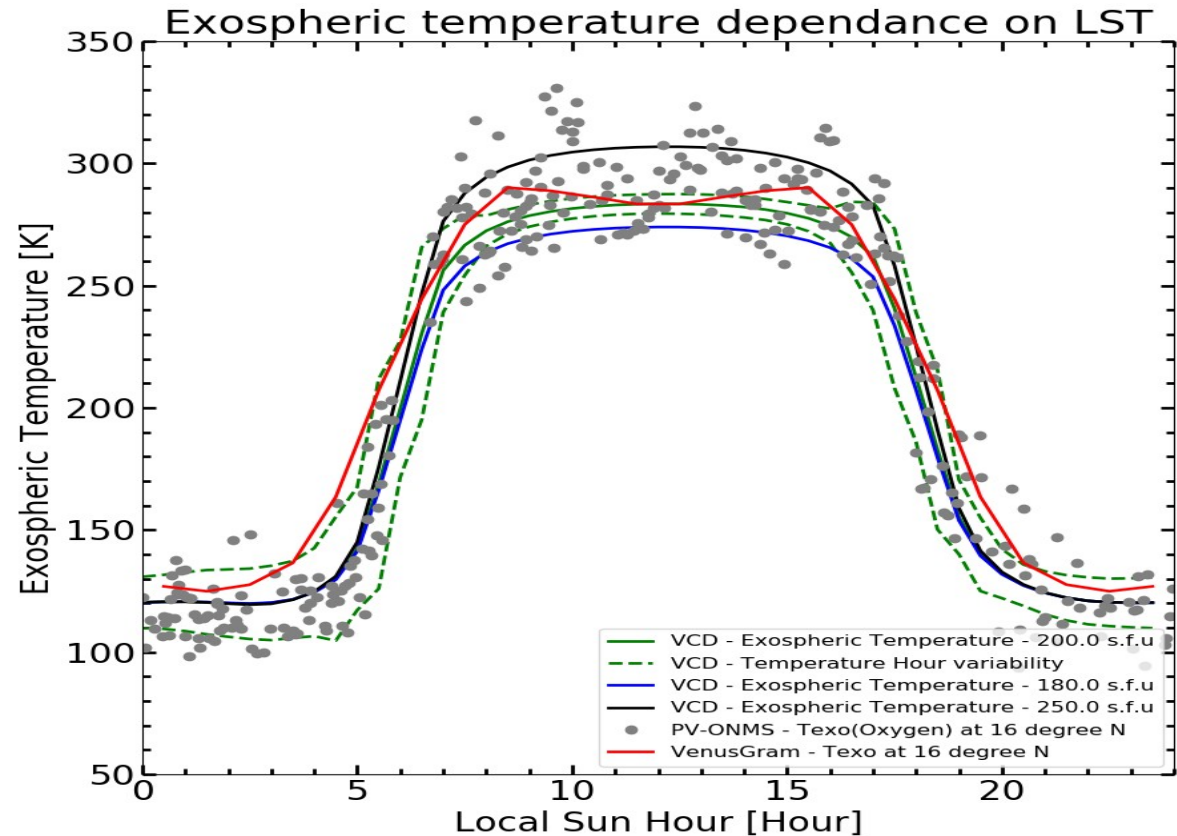
Temperature in the deep atmosphere



Comparison of potential temperature profiles between some Pioneer-Venus probes (day/night/North), the VCD (standard albedo scenario +RMS) and VenusGram2005

Temperature above 150 km

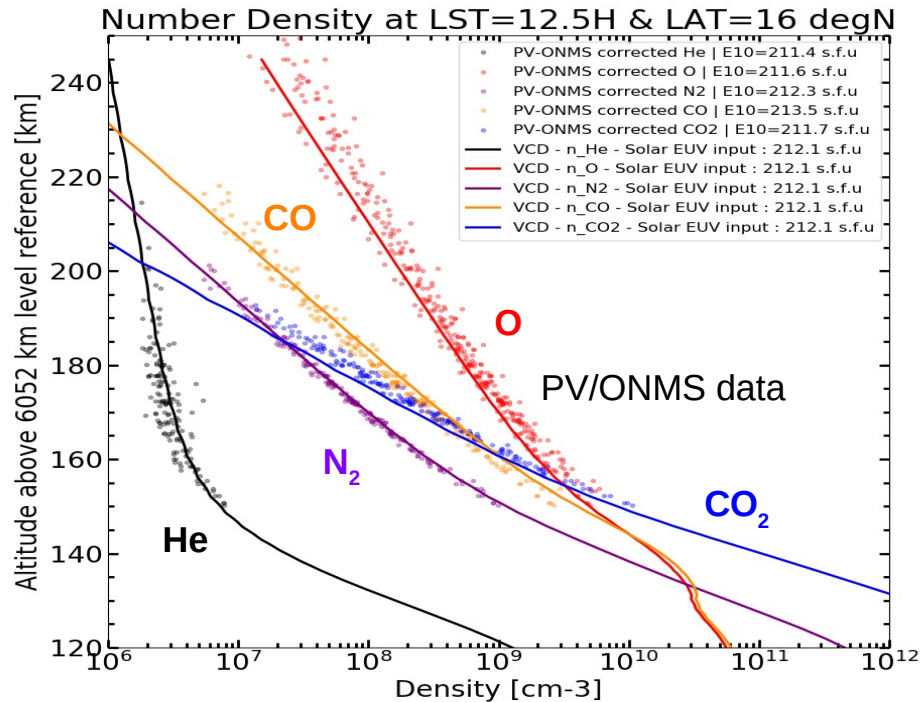
Exospheric temperature retrieved from O profiles in PV-ONMS datasets



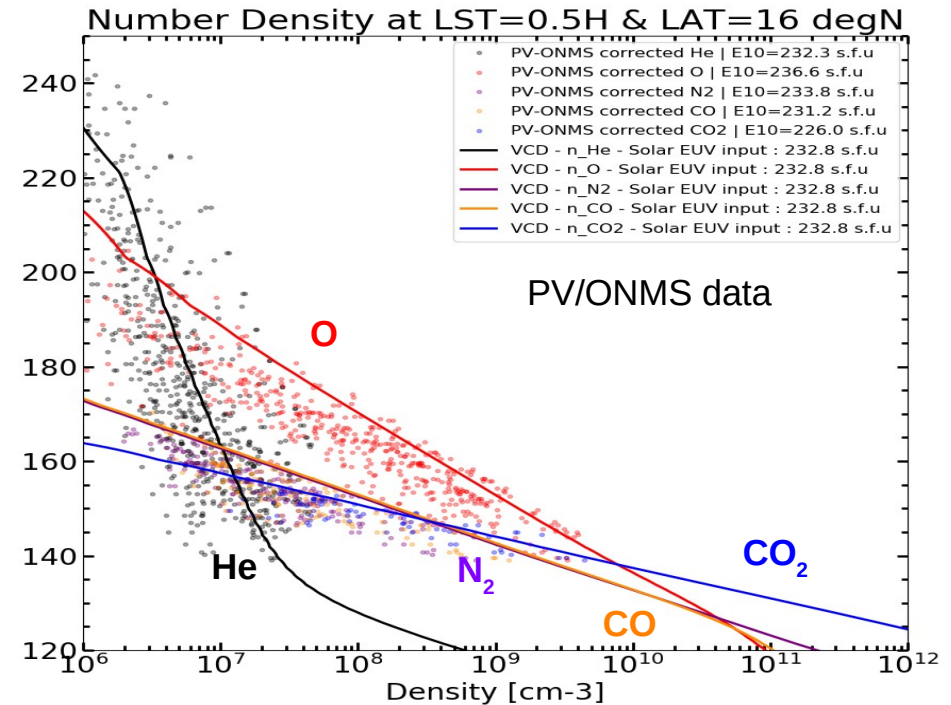
Composition: O and CO

O plays a significant role on temperature !
Increasing CO₂ photodissociation improves O and CO...
Investigations ongoing

Noon

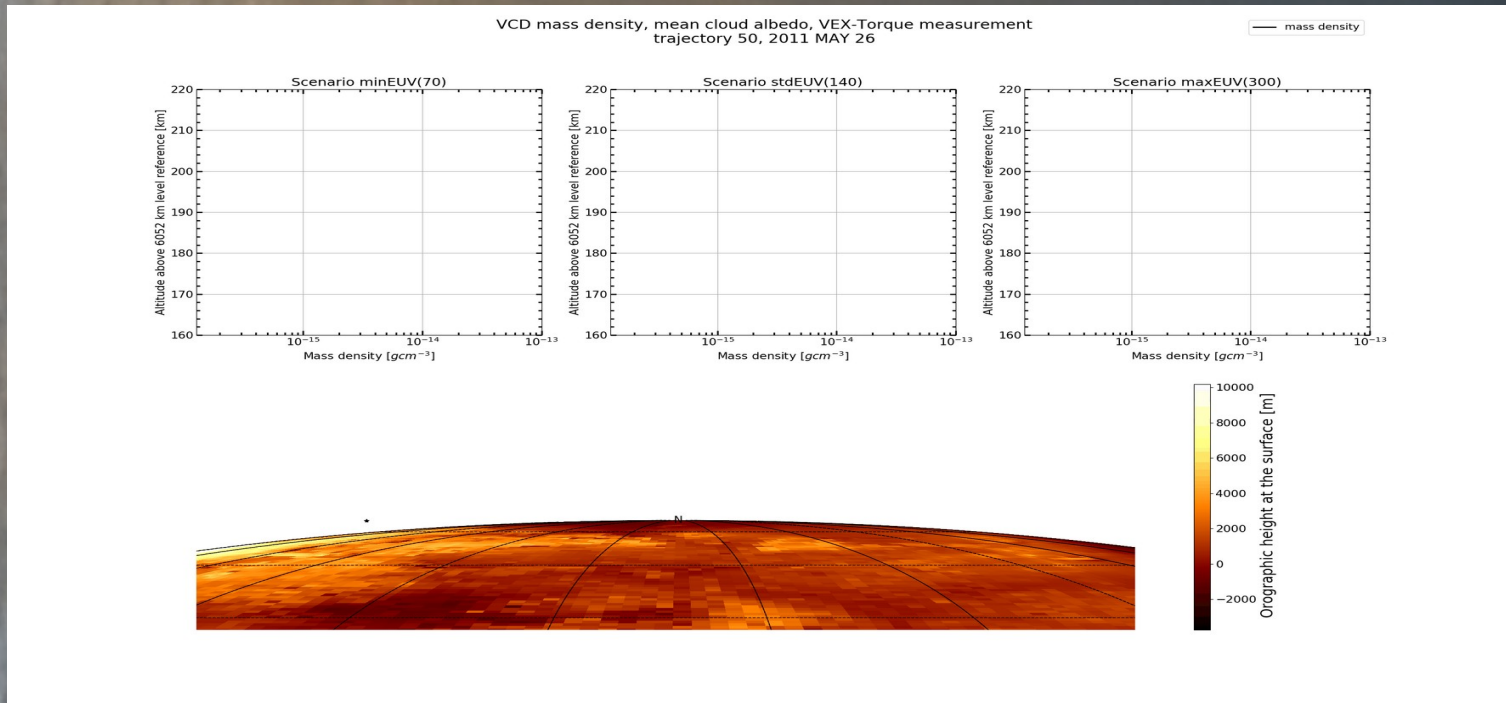


Midnight



Examples of applications

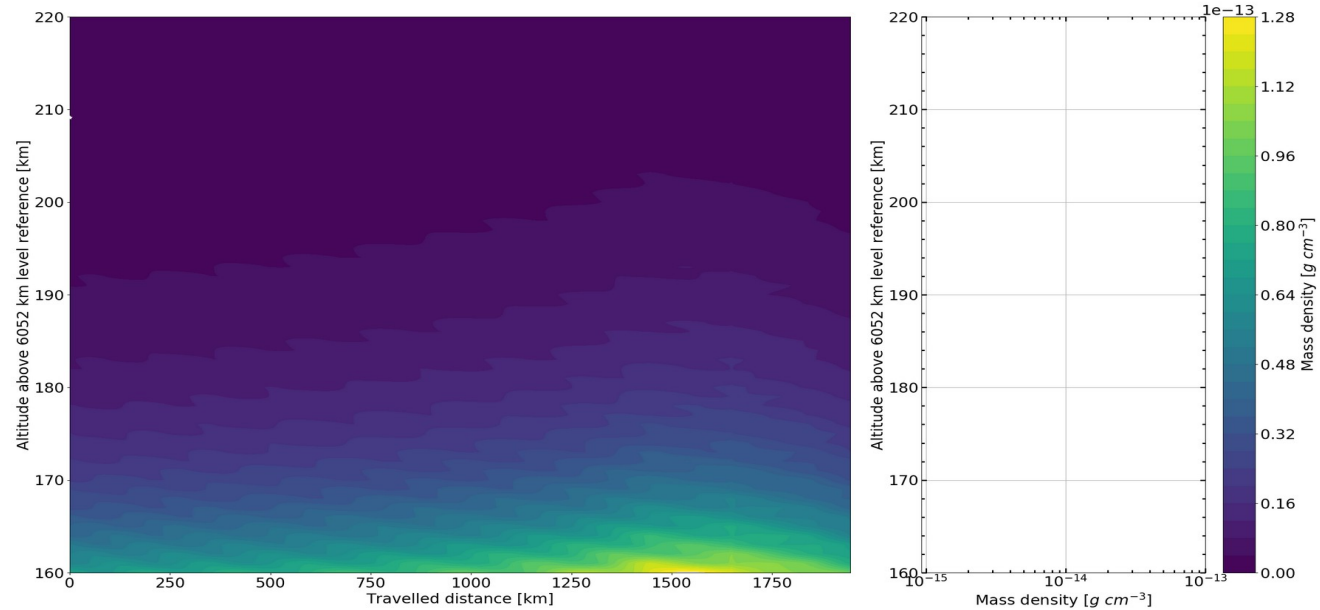
Illustrative aerobraking example



Examples of applications

Aerobraking example

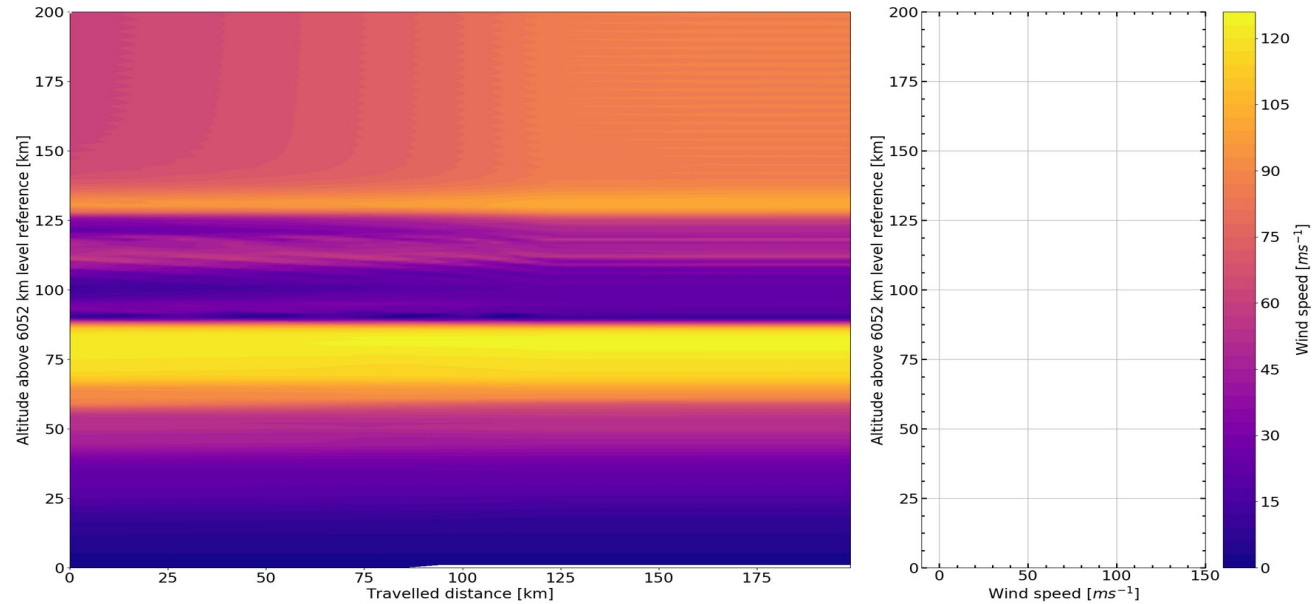
VCD density, stdEUV(140), mean cloud albedo, VEX-Torque measurement
trajectory 50 from (75 N,9 W) to (87 N,170 E), 2011 MAY 26



Examples of applications

Descent probe example

VCD wind speed, small and large scale perturbations, stdEUV(140), mean cloud albedo, DAVINCI trajectory 0
from (13 S,1 E) to (24 S,2 E), 2028 APR 10



Using the VCD

The VCD access software is provided as a Fortran subroutine `call_vcd` which provides point-wise query that one should use in their own home-made codes

```
call call_vcd(z_key,z,lon,lat,hires_key,
              date_key,juliandate,
              localtime,dset,EUV_scena,
              albedo_scena,varE107,
              perturb_key,perturb_seed,
              perturb_gw_length,extvar_keys,
              zon_wind,mer_wind,vert_wind,
              temp,pres,dens,extvar,
              seed_out,ier)
```

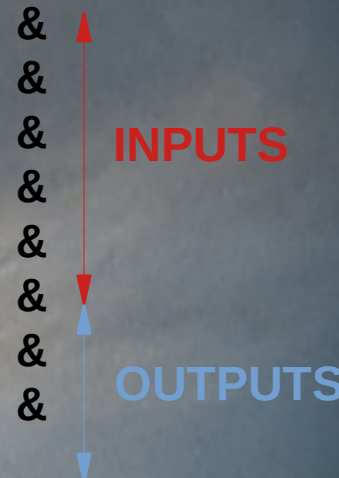


Diagram illustrating the inputs and outputs of the `call_vcd` subroutine:

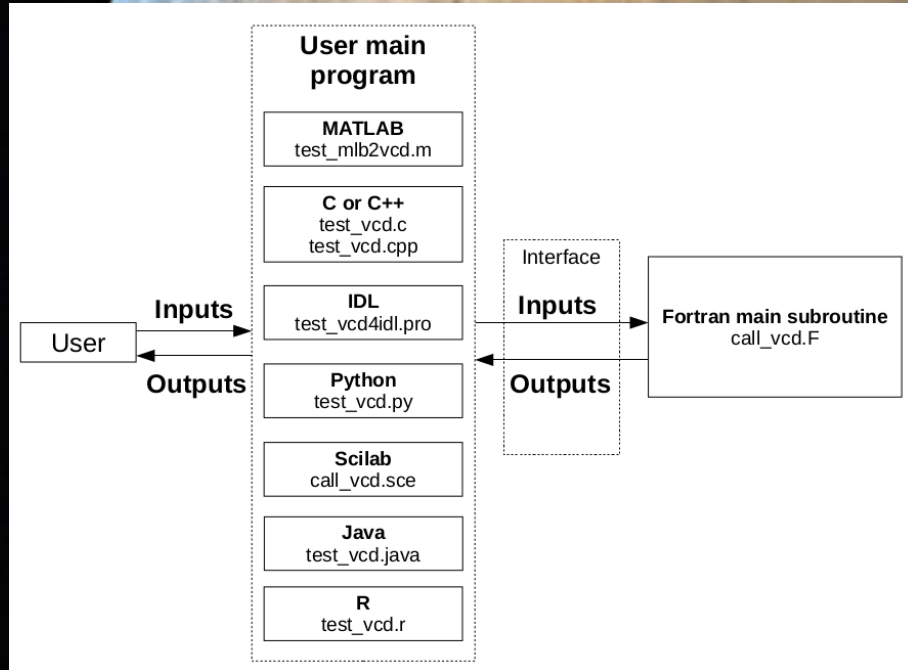
- INPUTS (Red arrow pointing up):** `&z_key`, `&juliandate`, `&localtime`, `&dset`, `&EUV_scena`, `&albedo_scena`.
- OUTPUTS (Blue arrow pointing down):** `&varE107`, `&perturb_seed`, `&extvar_keys`, `&zon_wind`, `&mer_wind`, `&vert_wind`, `&temp`, `&pres`, `&dens`, `&extvar`, `&seed_out`, `&ier`.

See the [VCD User Manual](#) for a very detailed description of all inputs and output arguments

Various interfaces for the VCD

A full version with various interfaces:
Fortran, C, C++, IDL, Matlab, Python...

A web interface for quick-look and plots:
<http://www-venus.lmd.jussieu.fr/>



Venus Climate Database v2.3 : The Web Interface

[Information] [Gallery] [Report issue] [Credit] [How to cite the VCD]

VARIABLE(S) TO DISPLAY OR DOWNLOAD	TIME COORDINATES	SPATIAL COORDINATES
Variable 1: Temperature (K) Variable 2: (None) Variable 3: (None) Variable 4: (None)	<p>VENUS date Local Time: 0 Venusian hour write a value (or) a range 'val1 val2' (or) 'all'</p> <p>EARTH date YY / MM / DD @ hh:mm:ss UTC 2023 / 8 / 3 @ 10:11:28 Use Earth date to calculate Venus Ls</p> <p>Earth Julian Date: 2460159.9247 Venus Solar Longitude: 66.9</p>	<p>write a value (or) a range 'val1 val2' (or) 'all'</p> <p>Latitude: all degree North Longitude: all degree East Altitude: 10. m above surface</p> <p>Spacecraft landing site & date: (None)</p>

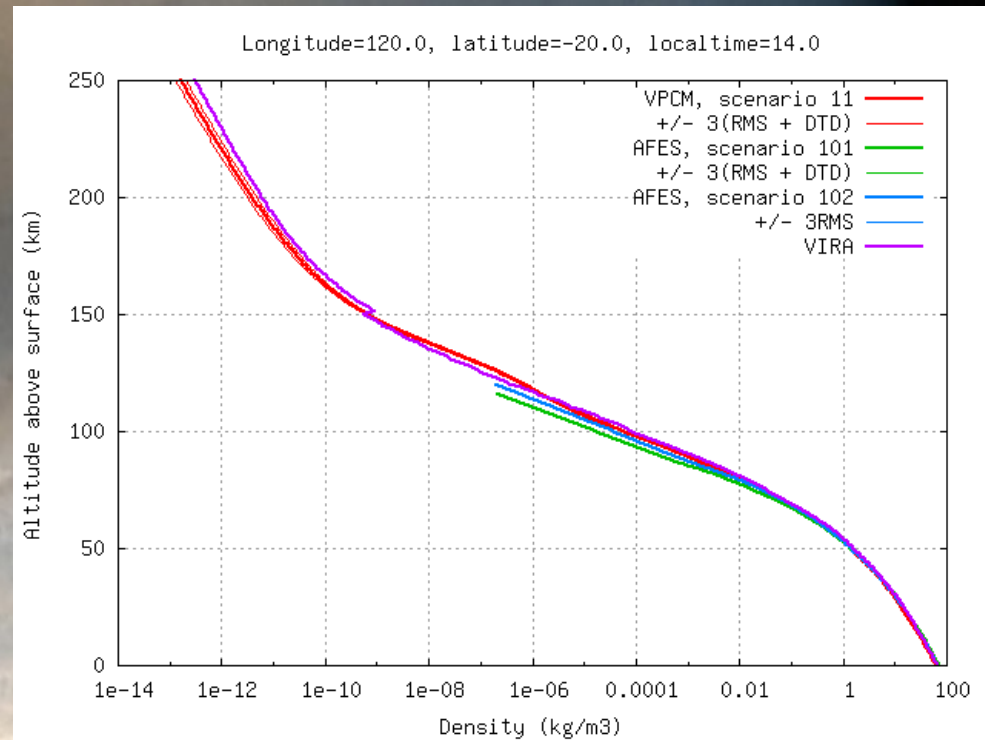
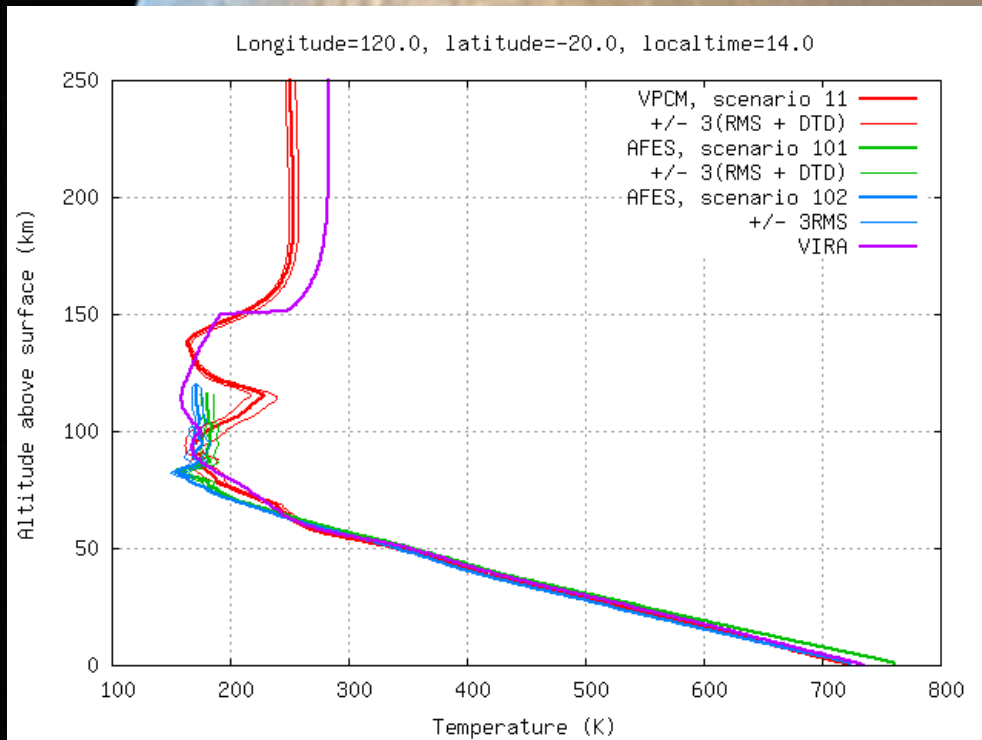
CUSTOMIZE DATA REQUEST	CUSTOMIZE FIGURES	
<p>Same localtime everywhere: off on</p> <p>Cloud albedo scenario: Standard cloud albedo</p> <p>Solar EUV scenario: Average solar EUV</p> <p>High-resolution topography: off on</p> <p>Averaging (only for 2D plots for now!): off zonal diurnal</p>	<p>Figure format: PNG PNG hi-res EPS</p> <p>[1D] Log(values): off on</p> <p>[2D] Colormap: blue green yellow red</p> <p>[2D] Values range: to</p> <p>[2D map] View: flat</p> <p>[2D map] Transparency (%):</p> <p>[2D map] Wind vectors: off on</p> <p>[2D map] Add marker at lat: lon</p>	<p>SUBMIT</p> <p>RESET</p>

Future plans: Moving on to VCD 3.0

- A request by ESA (and some VCD users) : to have the possibility of obtain from the VCD, a “**multi-GCM variability**” evaluation (based on the range of predictions variation due to the difference in the climate models themselves).
- We have thus proposed to the community to gather “reference simulations” to integrate in the next version of the VCD, along a pre-defined protocol, with some positive responses from:
 - AFES GCM Team (including some ALEDAS-V assimilation), Japan
 - VTGCM Team, USA
 - TUGCM Team, Tohoku University, Japan
 - ...
- To this day, only AFES & ALEDAS-V data has been obtained and processed, enabling the generation of a beta-version of VCD3.0. To be continued...

Illustrative example of VCD 3.0

Some profiles from VCD3.0 at same location and time but from the different models:



Illustrative example of VCD 3.0

Some profiles from VCD3.0 at same location and time but from the different models:

