



Venusian atmospheric general circulation model for the Earth Simulator

Local ensemble transform Kalman filter data assimilation system (VALEDAS)



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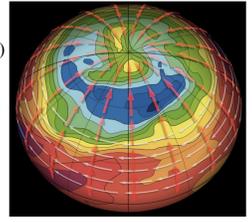
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1. Introduction

- AFES*-Venus (*Atmospheric GCM For the Earth Simulator)
- Dynamical Venus-GCM highly optimized for parallel vector super computer
- Start from idealized super-rotation: Saving computational cost with high resolution run
- Maintain super-rotation with realistic setting: under the realistic solar heating and static stability

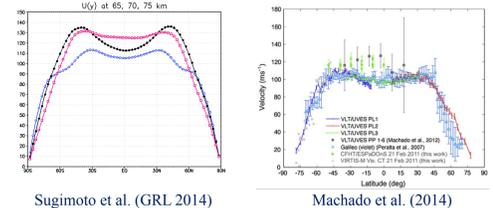
Resolutions from T42L60 to T639L260 (1920 × 960 grids with 260 layers)

- Focus on atmospheric motions near the cloud level
- Super rotation driven by mean meridional circulation (Sugimoto et al., GRL 2019)
- Baroclinic instability & Neutral waves (Sugimoto et al., JGR 2014; GRL 2014)
- Thermal tide & Short period waves (Takagi et al., JGR 2018; Ando et al., JGR 2018)
- Polar vortex & Cold collar (Ando et al., Nature Com.2016; JGR 2017)
- Energy spectra & Streak structures (Kashimura et al., Nature Com.2019)



Ando et al. (Nature com. 2016)

Final goal: To understand fundamental mechanism of super-rotation

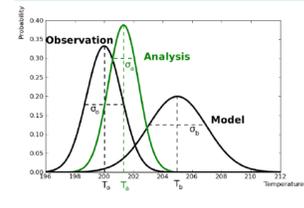


Sugimoto et al. (GRL 2014)

Machado et al. (2014)

- LETKF (Local Ensemble Transform Kalman Filter)
- Simple data assimilation method but computationally expensive
- LETKF applied to Earth and Mars: Hunt et al.(2007), Miyoshi et al.(2007...), Hoffman et al.(2010)

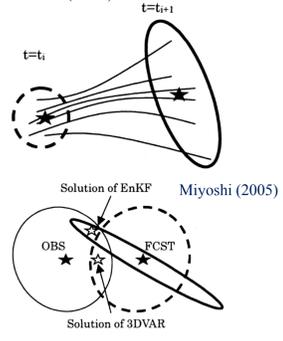
Local: considers only observations within a certain distance.
Ensemble: uses an ensemble of GCM forecasts.
Transform: uses a square-root filter.
Kalman Filter: uses past information to update the present state, and estimates both the state and its uncertainty (covariance)



$$\frac{1}{\sigma_a^2} = \frac{1}{\sigma_o^2} + \frac{1}{\sigma_b^2}$$

$$T_a = T_b + \frac{\sigma_b^2}{\sigma_o^2 + \sigma_b^2} (T_o - T_b)$$

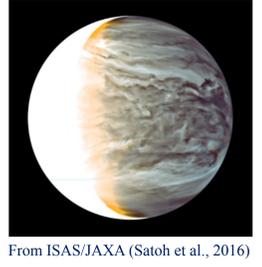
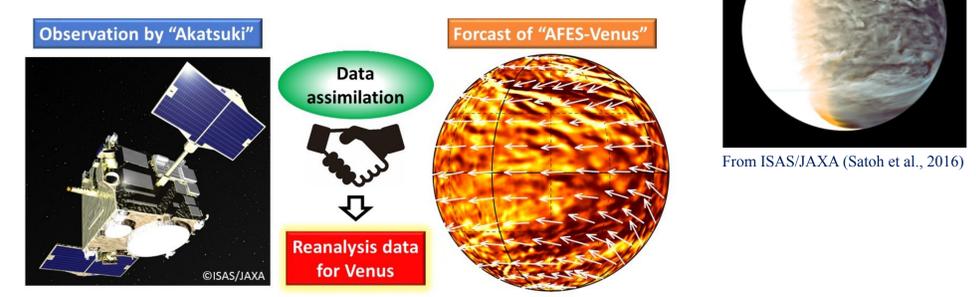
From Dr. T. Navarro (personal com.)



Miyoshi (2005)

AFES-LETKF data assimilation system provides Re-Analysis (ALERA) data for the Earth

- Akatsuki Venus Climate Orbiter
- Frequent observations with multiple altitude: Nakamura et al. (2011, 2014...)
- Datasets will be available for the data assimilation



From ISAS/JAXA (Satoh et al., 2016)

This study: AFES-LETKF data assimilation system for the Venusian atmosphere

2. Experimental setting

- AFES-Venus (VAFES)
 - 3-D Primitive equation on sphere (hydro static balance) without moist processes
 - Resolution: T42L60 (128 × 64 × 60)
 - Specific heat: Cp is constant (1000 Jkg⁻¹k⁻¹)
 - Horizontal hyper-viscosity: 0.1 Earth days for 1/e
 - Vertical eddy viscosity: 0.15 m²s⁻¹
 - Rayleigh friction: lowest and above 80 km (sponge layer except for zonal flow)
 - No topography and planetary boundary layer

AFES-Venus runs	
Cases	Solar heating
Qz	Zonal mean component only
Qt	Including diurnal variation

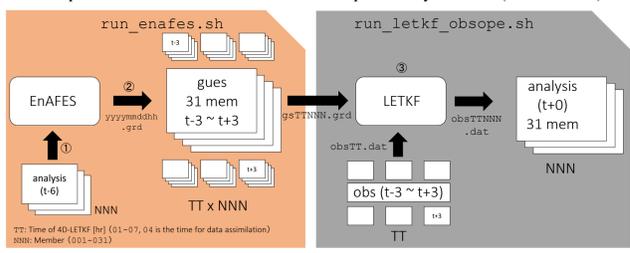
- Solar heating
 - Zonal (Qz) and diurnal (Qt) component of realistic heating
 - Based on Tomasko et al. (1980) and Crisp (1986)
- Infrared radiative process
 - Simplified by Newtonian cooling: dT/dt = -κ(T - T_{ref}(z))
 - κ: based on Crisp (1986)
 - T_{ref}(z): horizontally uniform field

Test observations			
Case	Obs	interval	AFES
H1	Qt	1h	Qz
Vqz	VMC	~24h	Qz
Vqt	VMC	~24h	Qt

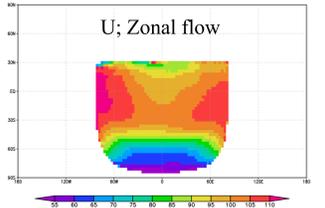
- Test observations; horizontal flows at the cloud top level
 - Idealized: AFES-Venus runs; Qt (including diurnal variation)
 - Real: VMC; Venus Monitoring Camera (73 obs. in Epoch 4: 28 Jan to 26 Apr 2008)

- AFES-Venus runs; Qz and Qt
 - T42L60: 31-member ensemble, 6-hourly assimilation cycle interval
 - Localization: horizontally 400 km, vertically lnP~0.4
 - Observational errors: 4.0 m/s
 - Inflation: 10 %

- Flow chart of the VAFES-LETKF data assimilation system
 - 9-hour forecast from t=0 and use from t=3 to t=9 for the assimilation
 - Input observations from t=3 to t=9 and output reanalysis at t=6 (=4D LETKF)



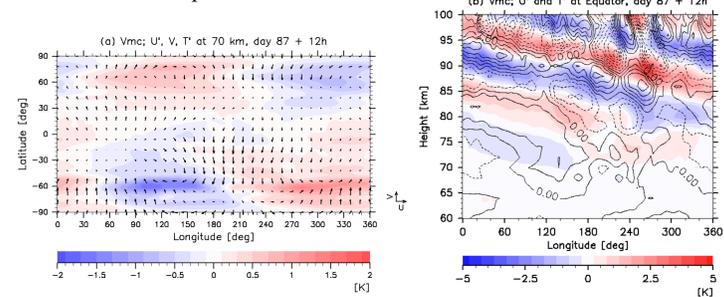
Structures of the thermal tide in VMC



Kouyama et al. (JGR 2013)

3. Results

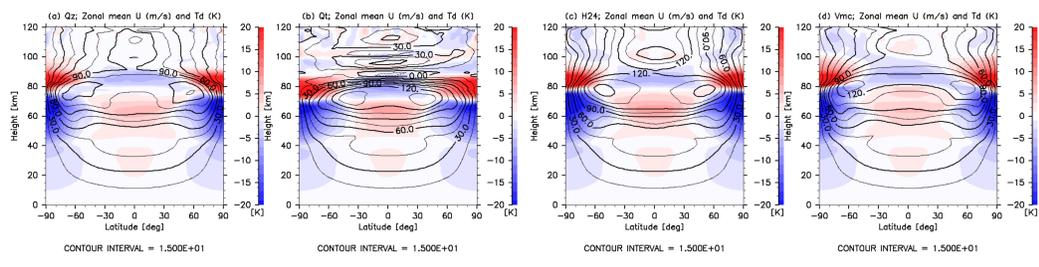
- Structures for Case Vqz



Thermal tides appear but amplitudes of them are very weak

- Latitude-height cross sections for Cases Qz, Qt, H1, Vqz

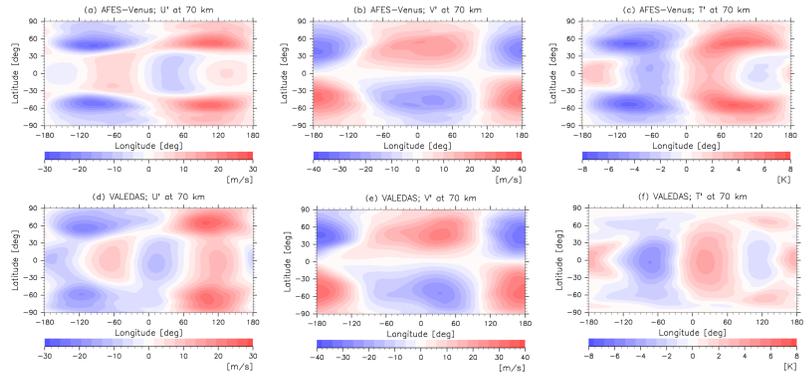
Sugimoto et al., Sci. Rep. 2017



U: zonal mean zonal flow (contour), Td: temperature deviations from the horizontally averaged temperature (colour)

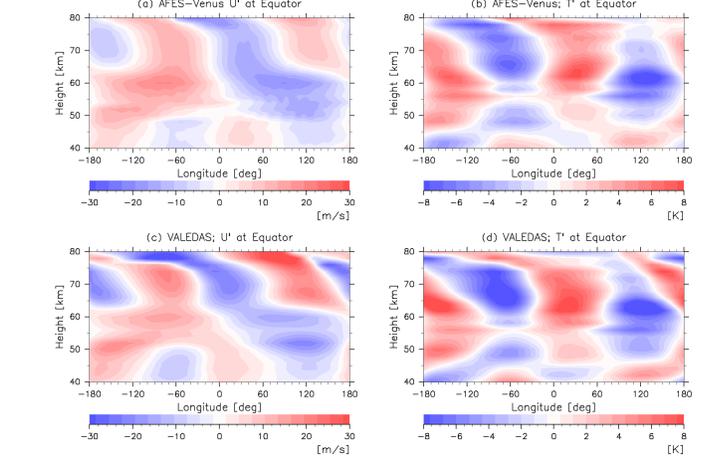
General circulation can not be changed by the data assimilation with sparse observations.

- Horizontal structures for Cases Qt, Vqt



Phases of thermal tides are improved not only horizontal winds but temperature.

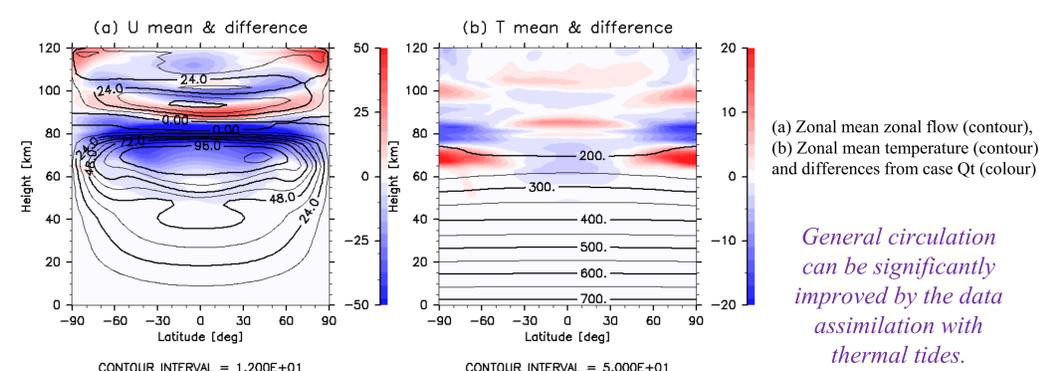
- Vertical structures for Cases Qt, Vqt



Vertical structures of thermal tides are also improved.

- Latitude-height cross sections for Case Vqt

Sugimoto et al. GRL, 2019



General circulation can be significantly improved by the data assimilation with thermal tides.

4. Summary

- AFES-LETKF data assimilation system for the Venusian atmosphere has been developed.
- VMC/VEX horizontal winds improve phases of thermal tides and general circulation.
- Akatsuki frequent observations will improve AFES-Venus forecast significantly.

References:

Sugimoto, N., A. Yamazaki, T. Kouyama, H. Kashimura, T. Enomoto and M. Takagi, (2017), "Development of an ensemble Kalman filter data assimilation system for the Venusian atmosphere", Scientific Reports, 7, 9321, 9pp, doi:10.1038/s41598-017-09461-1.

Sugimoto, N., T. Kouyama and M. Takagi, (2019), "Impact of data assimilation on thermal tides in the case of Venus Express wind observation", Geophys. Res. Lett., 46, doi:10.1029/2019GL082700.

See P29 also "The study on the reproducibility of cold collar assuming radio occultation measurement by small satellites"