

# Venus atmosphere dynamics revealed by cloud tracking using images from Akatsuki

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### Cloud tracking method → poster P39

- A novel automated method: a key point
  - Based on the cross-correlation method but using 3 or more images (Ikegawa and Horinouchi 2016: IH16)
  - Sophisticated **error correction** and **quality control** (Horinouchi et al., 2017a, IH16)

→ high-quality gridded horizontal wind dataset (much greater # of wind vectors than manual tracking)

## Equatorial jet in the lower and middle cloud regions (Horinouchi et al 2017b)

- Used nightside images from the 2-μm camera IR2
  - Near-infrared "window"
  - Visualizes the opacity contrasts mainly in the lower and middle clouds (45-60 km)
  - Analysis period: Mar & Jul-Aug, 2016







- 2 hourly 2.26µm radiance
- (2016-07-11 12; orthographic map projection)
- Moves westward: superrotation
- White-out: from dayside (not in the deonvolved version: Poster by Satoh)

### Zonal flow: exhibits a jet-like feature at low latitude: "equatorial jet" (observed through July and Aug, but not in Mar)



Cloud top wind from UV images (Akatsuki UVI)  $\rightarrow$  no equatorial jet

#### Previous studies using near-IR nightside images did not detect an equatorial jet. --- because time sampling was limited?



IR2 nightside images exhibit rich dynamical features: e.g. Peralta et al, 2018,9 Limaye et al 2018

### A feature like vortices due to barotropic/baroclinic instability (2016-08-25)



### Really rolling-up? – Yes, likely.

 arrows: deviation from the solid-body rotation of -75 m/s at the equator (CAUTION: based on previous version of geographical mapping)



Wind estimation will be updated by further refining geographical mapping



Cloud top mean winds and shear (updated from Horinouchi et al 2018 E.P.S.)

- Merit of Akatsuki: can observe both hemispheres equally, etc.
- We used dayside images from the UV camera UVI
  - 365 nm (traditional, unknown UV absorber)
  - 283 nm (SO2; novel limited obs by PVO)
  - Analysis period: Dec 2015 Dec 2018



#### UV Imager on Akatsuki



### Example 2015-12-07 365nm (deviation from a solid-body rotation: 90 m/s at EQ)



cmv\_radiance.rb --hipass 4 --eps 10 --rmax 0.6 --latran -30..30 --lonran 70..... 2017-04-26

2017-04-26

Example 2016-05-06 365nm (deviation from a solid-body rotation: 90 m/s at EQ)

#### $[U+90\cos\varphi, V]$ & radiance



2017-04-12

cmv\_radiance.rb -- hipass 4 -- eps 10 -- rmax 0.6 -- latran -30..30 -- lonran 180....







ltm\_winds\_lines.rb --eps\_errbar cmv\_v20190201\_nt3ll/daily\_uvi\_365\_eps20\_pp0hvlen2.0\_LT.nc --tstr 2015-12-07,2016-09-21,2017-04-21,2017-12-01,2018-06-15 --tend 2016-09-2... 2019-04-14



ltm\_winds\_lines.rb --eps\_errbar cmv\_v20190201\_nt3ll/daily\_uvi\_283\_eps20\_pp0hvlen2.0\_LT.nc --tstr 2015-12-07,2016-09-21,2017-04-21,2017-12-01,2018-06-15 --tend 2016-09-2... 2019-04-14

# Cloud top angular momentum & heat fluxes (submitted)

- Previous studies
  - Momentum flux: Rossow et al (1990), Limaye (2007) → inconclusive
  - Heat flux: None
- Here we used
  - Winds from UVI 365 nm images
  - $\sim$  Cloud-top temperature from LIR (obs at high emission angles)

#### LIR



UV Imager on Akatsuki





### Mom flux by thermal tide: accelerating super-rotation at low lat.

### Mom flux by transient disturbances: decelerating super-rotation (more weakly)



### Meridional heat flux by zonal wavenumber 1 tide



Vertical component of the EP flux associated with the tidal heat flux

0

Suggested angular-momentum balance and meridional circulation in the cloud layer of Venus



 $\rho_0 m_t + \boldsymbol{\nabla} \cdot \rho_0 m(\boldsymbol{v}^*, \boldsymbol{w}^*) = \boldsymbol{\nabla} \cdot \boldsymbol{F}$ 

F: EP flux (wavy arrows)

(v\*,w\*) : "residual circulation" (mean Lagrangian meridional circulation including the Stokes drift by eddies)

Suggested overall circulation in the Venusian cloud layer is Earth-like

### Conclusions

- Cloud tracking using data from Akatsuki provided new knowledge on the Venusian atmosphere
  - From the nightside through-cloud imaging
    - Variability of flows in the lower-to-middle cloud layers: e.g. equatorial jet (formed tentatively), unstable vortices,…
  - From cloud top winds
    - Mean vertical shear near the cloud top
    - Meridional asymmetry (which may be due to the asymmetry in the vertical distribution of UV absorbing substances)
  - From cloud top winds and temperature
    - Meridional angular momentum flux and heat flux
    - Providing insights on the meridional circulation and angular momentum balance
       → super-rotation maintenance processes
- Further observations and studies will elucidate more and provide deeper understanding of the Venusian atmosphere