High-pressure physics of the Earth and beyond













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High-P,T phase relation of Earth materials (ex. SiO₂) Hugoniot (Hicks+ 2006 PRB)



- Experimental investigations currently almost impossible
 - *Ab initio* theoretical computation method

Ab Initio (first principles) Earth and Planetary Sciences

(i) Structural exploration

--- Shear response





1.Fundamental methodologies of the ab initio electronic structure calculation method

2.Applications to high-pressure mineral physics and Earth & planetary interiors

- Phase relations including melting
- Electronic property
- Transport property



Schrödinger equation

$$\widehat{H}(\mathbf{r}_1,\mathbf{r}_2\cdots)\Psi(\mathbf{r}_1,\mathbf{r}_2\cdots)=E\Psi(\mathbf{r}_1,\mathbf{r}_2\cdots)$$

 \widehat{H} : Hamilton operator (Hamiltonian) Ψ : Wave function (Eigen vector) E: Total energy (Eigen value)

Eigenvalue problem

Quantization

$$\hat{p} = -i\hbar \nabla$$
 (Momentum)
 $\hat{E} = i\hbar \partial / \partial t$ (Energy)

Many-electron system

Interacting electrons

One electron in an effective potential





$$\begin{bmatrix} -\frac{\hbar^2}{2m} \Delta + V_n[n(\mathbf{r})] + V_H[n(\mathbf{r})] + V_{XC}[n(\mathbf{r})] \end{bmatrix} \phi_i(\mathbf{r}) = \varepsilon_i \phi_i(\mathbf{r}),$$

$$n(\mathbf{r}) = \sum_i |\phi_i(\mathbf{r})|^2 \qquad \qquad \text{Kohn-Sham equations (DFT)}_{(Hohenberg \& Kohn, 1964; Kohn \& Sham, 1965)}$$

Density Functional Theory (DFT)

One-electron Hamiltonian

$$n(\mathbf{r}) = \sum_{i} \left| \phi_{i}(\mathbf{r}) \right|^{2}$$

Angular component of a wavefunction in a central force field (Coulomb potential) $\propto 1/r$ = Spherical harmonics $\Phi_{l.m}$

l: angular momentum quantum number (s, p, d, f) *m*: magnetic quantum number

+ spin quantum number (up or down)



XC (exchange-correlation) potential ($V_{\rm XC}$)

Electron (fermion) → Quantum many-body effects



Local density approximation (LDA)

 $V_{\rm XC}$ determined for the homogeneous electron gas, which can be calculated precisely, is applied also to general systems.

Quantum Monte-Carlo

Energy level to energy band



Electronic density of states

T Tsuchiya, 9th ISPS, 25 June 2012 Tsuchiya & Tsuchiya (2011) PNAS





B2-CaO 0.4 TPa



Pseudo-potential approximation

- Valence electrons only contribute chemical bonding.
- Nuclei + Core electrons → Ion potential (with orthogonality of valence and core electrons)



PP is determined nonempirically to reproduce the true wavefunction correctly in the bonding region ($r > r_c$).

Self-Consistent Field (SCF) method

$$\hat{h}_i[n(\mathbf{r})] \phi_i(\mathbf{r}) = \varepsilon_i \phi_i(\mathbf{r})$$

 ϕ_i depends on the Hamiltonian $\hat{h}_i[n(\mathbf{r})]$.

 $\hat{h}_i[n(\mathbf{r})]$ depends on the total charge density $n(\mathbf{r})$.

 $n(\mathbf{r})$ depends on the wave functions ϕ_i .

The equation is solved iteratively.

Valence charge density of some representative bond types





Large scale computation

GRC-SRFC parallel clusters









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http://jp.fujitsu.com

Atomic dynamics and temperature effects $DFT \rightarrow Total energy of many-electrons system (E_{tot})$ + Hellman-Feynman theorem (perturbation theory)



 $\frac{\partial^2 E_{tot}}{\partial \mathbf{r}_i \partial \mathbf{r}_j} \text{ (Force constant)} \rightarrow Ab \text{ initio} \text{ lattice dynamics} \text{ (LD)}$

Molecular Dynamics method

A method to investigate dynamical property of many-atom systems



Time evolution is calculated by numerically integrating the Newton's equation of motion

$$\mathbf{x}_{i}(t + \Delta t) = \mathbf{x}_{i}(t) + \mathbf{v}_{i}(t)\Delta t + \frac{1}{2}\mathbf{a}_{i}(t)\Delta t^{2}$$
$$\mathbf{v}_{i}(t + \Delta t) = \mathbf{v}_{i}(t) + \frac{\mathbf{a}_{i}(t) + \mathbf{a}_{i}(t + \Delta t)}{2}\Delta t \qquad \text{Velocity-Verlet}$$
algorithm

 $\mathbf{a}_i(t) = \mathbf{F}_i(t)/m_i$

Hellman-Feynman theorem

$$\frac{dE}{d\lambda} = \left\langle \psi(\lambda) \left| \frac{d\widehat{H}}{d\lambda} \right| \psi(\lambda) \right\rangle$$

Force acting on an atom can be calculated directly from the SCF charge density.



Macroscopic thermodynamic quantities

Temperature (Energy equipartition principle)

$$T = \frac{1}{3Nk_B} \sum_j m_j v_j^2$$

Pressure (Virial theorem)

$$P = P_{static} + \frac{Nk_BT}{V} - \frac{1}{3V} \sum_{i>j} \mathbf{F}_{ij} \otimes \mathbf{r}_{ij}$$

But not temperature dependence of heat capacity due to the classical Newton's dynamics (Dulong-Petit law)



Lattice Dynamics Method



Equation of motion

$$F^{\alpha}[\mathbf{r}_{n}(l)] = -\frac{\partial E^{harm}}{\partial u_{n}^{\alpha}(l)} = -\sum_{l',ll',\alpha\beta} \Phi_{nn'}^{\alpha\beta}(l,l')u_{n'}^{\beta}(l')$$

Solution $u_n^{\alpha}(l) = u_n^{\alpha}(\mathbf{q}) \exp[i\mathbf{q} \cdot \mathbf{r}(l) - i\omega t]$

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Dynamical matrix
$$\mathbf{D}_{nn'}^{\alpha\beta}(\mathbf{q}) = \frac{1}{\sqrt{m_n m_{n'}}} \sum_l \mathbf{\Phi}_{nn'}^{\alpha\beta}(0,l) \exp[i\mathbf{q} \cdot (\mathbf{r}_0 - \mathbf{r}_l)]$$



$$\omega^2 u_n^{\alpha}(\mathbf{q}) = \sum_{n',\beta} \frac{\mathbf{D}_{nn'}^{\alpha\beta}(\mathbf{q})}{\sqrt{m_n m_{n'}}} u_{n'}^{\beta}(\mathbf{q})$$

Density Functional Perturbation Theory (DFPT) (Baroni+ PRB 1987; RMP 2001)



Phonon dispersion of MgSiO₃ pPv (P=120GPa)

Calculate the dynamical matrix based on the quantum perturbation theory

Phonon (quantized lattice vibration) **dispersion relation**

gold (Au) 300 200 ω (cm⁻¹) 100 Х κ Г Г q-vector Calc $V/V_0 = 1$ 0.93 0.86 0.82 0.79 Expr. $(V/V_0=1)$ Ο (Neutron Scatering)

Tsuchiya (2003) JGR

Silicates and solid solutions

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(Mg,Fe)SiO₃ Metsue & Tsuchiya (2011,2012)

Quasi-Harmonic Approximation (QHA)

Phonon Helmholz free energy

$$F_{ph}(V,T) = \frac{1}{2} \sum_{\mathbf{q},j} h\omega_j(\mathbf{q},V) + k_B T \sum_{\mathbf{q},j} \ln\left[1 - \exp\left\{-\frac{h\omega_j(\mathbf{q},V)}{k_B T}\right\}\right]$$

Total Helmholz free energy

$$F(V,T) = U_{stat}(V) + F_{ph}(V,T) + F_{el}(V,T) + \cdots$$

Pressure
$$P = -\left[\frac{\partial F}{\partial V}\right]_T$$
 Entropy $S = -\left[\frac{\partial F}{\partial T}\right]_V$

Other thermodynamic functions including

$$G(P,T) = F(V,T) + P(V,T)V$$

Crystal thermodynamics (e.g., MgSiO₃ Pv and PPv)



Tsuchiya+ (2005) JGR

Akaogi+ (2008) Phys Chem Miner



Excellent agreement

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How accurate are calculated EoS? e.g., EoS of H₂O

Experimental confirmation in the TPa condition is quite difficult. Laser or magnetic shock technique seems hopeful.



Thermodynamic phase stability

Gibbs free energy

$$G(P,T) = F(V,T) + P(V,T)V$$



High-*P*,*T* phase boundaries can be determined.

Multicomponent (*P-V-T-x*) phase equilibrium



Tsuchiya & Tsuchiya (2008) PNAS

Internally consistent LDA+U for Fe-bearing system

$$E^{LDA+U}[n(\mathbf{r})] = E^{LDA}[n(\mathbf{r})] + E^{Hub}[\{n_m^{I\sigma}\}] - E^{DC}[\{n^{I\sigma}\}]$$

$$E^{Hub}[\{n_{mm'}^{I\sigma}\}] - E^{DC}[\{n^{I\sigma}\}] = \frac{U}{2} \sum_{I,\sigma} \operatorname{Tr}[\mathbf{n}^{I\sigma}(1-\mathbf{n}^{I\sigma})]$$

Spin transition in ferropericlase (Mg,Fe)O

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On-site Coulomb U parameter determined non-empirically based on a linear response formalism



Tsuchiya+ (2006) PRL



Crystal elasticity





 $\begin{pmatrix} \sigma_{11} & \sigma_{12} & \sigma_{13} \\ \sigma_{21} & \sigma_{22} & \sigma_{23} \\ \sigma_{31} & \sigma_{32} & \sigma_{33} \end{pmatrix}$ $\sigma_{ij} = V$

Strain ε_{ij}



$$\varepsilon_{ij} = \begin{pmatrix} \varepsilon_{11} & \varepsilon_{12} & \varepsilon_{13} \\ \varepsilon_{21} & \varepsilon_{22} & \varepsilon_{23} \\ \varepsilon_{31} & \varepsilon_{32} & \varepsilon_{33} \end{pmatrix}$$

$$\varepsilon_{ii} = \frac{\partial u_i}{\partial x_j}, \varepsilon_{ij} = \varepsilon_{ji} = \frac{1}{2} \left(\frac{\partial u_1}{\partial x_2} + \frac{\partial u_2}{\partial x_1} \right)$$

Elastic constant tensor

Linear response between stress and strain (Hooke's law)

$$\sigma_{ij} = \sum_{kl} c_{ijkl} \varepsilon_{kl}$$

 $\varepsilon_{ij} = \sum_{kl} s_{ijkl} \sigma_{kl}$

$$(i, j, k, l = 1 \sim 3)$$

or

Voigt notation (simplified notation)

$$11 \rightarrow 1, 22 \rightarrow 2, 33 \rightarrow 3, 23 = 32 \rightarrow 4, 31 = 13 \rightarrow 5, 12 = 21 \rightarrow 6$$

e.g., $c_{1111} = c_{11}, \quad c_{1122} = c_{12}, \quad c_{2323} = c_{44}$


Acoustic (elastic) wave speed

Equation of motion

$$\rho \frac{\partial^2 u_1}{\partial t^2} = \frac{\partial \sigma_{11}}{\partial x_1} + \frac{\partial \sigma_{12}}{\partial x_1} + \frac{\partial \sigma_{13}}{\partial x_1},$$

$$\rho \frac{\partial^2 u_2}{\partial t^2} = \frac{\partial \sigma_{22}}{\partial x_2} + \frac{\partial \sigma_{23}}{\partial x_2} + \frac{\partial \sigma_{21}}{\partial x_2},$$

$$\rho \frac{\partial^2 u_3}{\partial t^2} = \frac{\partial \sigma_{33}}{\partial x_3} + \frac{\partial \sigma_{31}}{\partial x_3} + \frac{\partial \sigma_{32}}{\partial x_3}$$

Orientational average

 \Rightarrow Polycrystalline wave speed

$$V_P = \sqrt{\frac{B_S + 4/3\mu}{\rho}}$$
 $V_S = \sqrt{\frac{\mu}{\rho}}$ $V_{\Phi} = \sqrt{\frac{B_S}{\rho}}$

Direct comparison with seismological observations

Cristoffel's equation

 $\left|c_{ijkl}\mathbf{n}_{j}\mathbf{n}_{l}-\rho V\delta_{ik}\right|=0$

⇒Single-crystal wave speed

e.g., ice X phase (P=130 GPa)



Shear response of crystals

Generalized Stacking Faults (GSF) Theory (Vitek, 1968; Cordier+ 2004; etc)



1.Fundamental methodologies of the ab initio electronic structure calculation method

2.Applications to high-pressure mineral physics and Earth & planetary interiors

- Phase relations including melting
- Electronic property
- Transport property

T Tsuchiya, 9th ISPS, 25 June 2012 **Example:** Post-perovskite transition in MgSiO₃

Upper mantle: Olivine Mg₂SiO₄



Upper mantle Transition zone Subduction Plume Lower mantle D"=150±100km **Outer core** CMB **Inner** core 364 23.513.5 0 [Pressure (GPa)] 136 6370 2890 660 410 0 [Depth (km)]



Lower mantle: Perovskite MgSiO₃

Perovskite to post-perovskite structural change





Tsuchiya+ (2004) EPSL



Perovskite (Lower mantle)



Shear deformation (ϵ_6)



Post-Perovskite (D" layer)

CaIrO₃-type structure

Orthorhombic cell (*Cmcm*, *Z* = 4) with SiO₆ octahedra.
Those octahedra are connected with each other by sharing edges along the a direction. This is a major reason for the structure more stable than Pv.

> Exper: Murakami+ (2004) Theor: Tsuchiya+ (2004) etc







Effects of Fe and Al incorporation

Table 2. Logarithmic Derivatives of Velocities and Density With Respect to Lateral Variations in the Fe and Al Content at 100 GPa

	$\partial \ln V_{\mathbf{P}} / \partial X$		$\partial \ln V_{\rm S} / \partial X$		$\partial \ln V_{\Phi} / \partial X$		$\partial \ln \rho / \partial X$	
	Fe	Al	Fe	Al	Fe	Al	Fe	Al
pv	-0.140	-0.049	-0.220	-0.083	-0.098	-0.026	0.231	0.017
ppv	-0.156	-0.057	-0.236	-0.131	-0.099	-0.012	0.228	0.021

Tsuchiya & Tsuchiya (2006) GRL



Observations (PREM, Dziewonski & Anderson, 1981)

Calculated velocities and density of some model rocks along adiabatic geotherm (Tsuchiya PEPI 2011; Tsuchiya & Kawai, under review)

V_s is the most insensitive to the composition among these three models. (cf, Murakami+ 2012)





Modeled velocity structure

Kawai & Tsuchiya (2009) PNAS

About CMB heat flux



Extend to planetary interiors



360GPa 6000 K

Rocky (silicate) mantle + iron core





Gassy (H,He + rock+ice core?) ~10M

e.g., Guillot (1999) Science

Now, more than 700 exoplanets have been found. Some of them are terrestrial = **Super-Earths (SE)**.



http://exoplanet.eu/



GJ1214 (~6.55 M_{\oplus} rock+H₂O)







Rivera+ (2005); Udry+ (2007); Charbonneau+ (2009); Barnes+ (2010); etc

High-P,T phase relation of Earth materials (ex. SiO₂) Hugoniot (Hicks+ 2006 PRB)



- High-*P*,*T* phase relations highly unclear at ultrahigh pressures
- But several important advances made by ab initio calculations

Hydrogen

T Tsuchiya, 9th ISPS, 25 June 2012 *Morales+ (2010) PNAS* Ab initio MD



Water

Schwegler+ (2008) PNAS Ab initio melting curve



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Ice

Hermann+ (2012) PNAS Ab initio static phase stability



Carbon "Diamonds in the sky" Ross+ (1981) Nature ?

Correa+ (2006) PNAS Ab initio high-P,T phase diagram of C



Knudson+ (2008) Science Experimental confirmation of the Dia-BC8-Liq triple point



Iron

Morard+ (2011) HEDP Ab initio melting curve



No liquid iron core in giant planets

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Stixrude (2012) PRL Ab initio High-P,T phase diagram



High-*P*,*T* **phase relation of SiO**₂



Dense packing structures with AX₂ stoichiometry for initial models (15 in total)

(1) α -PbO₂ (VI)



(2) Pyrite (FeS₂) (VI)





(10) MoSi₂ (X) (6) $P2_1/m$ (VIII) (7) Cotunnite (PbCl₂) (8) $Fe_2P(IX)$ (9) P-1 (IX) (IX)



(14) MgZn₂ (XII)







Blue: Si Red: O

Coordination numbers in parentheses

Static enthalpy differences relative to the pyrite phase





- Pyrite directly transforms to the Fe₂P-type phase (not cotunnite) at ~7 Mbar!
- No other stable structures
- No elemental dissociation of SiO_2 into Si (hcp) plus O_2 (ζ), either

Tsuchiya & Tsuchiya (2011)

Crystal structure of the Fe₂P-type new high-P phase of SiO₂

Tsuchiya & Tsuchiya (2011) PNAS

Hexagonal cell (P62m, Z = 3) with two different kinds of SiO₉ tricapped trigonal prisms
Those polyhedra are fairly regular. This makes the structure more stable than Cot.

s_{12} s_{1} s_{1} s_{1} s_{2} s_{2} s_{2} s_{3} s_{2} s_{3} $s_{$



O2 V_{poly} = 9.295 Å³ ECoN = 8.5

Si2-O1 (×6) = 1.555 Å

-O2 (×3) = 1.707 Å

02

01

02

01





(1) Quartz SiO₄





SiO₄₊₆ (3) Li₂ZrF₆-type GEODYNAMOS RE (metastable) SiO₆



(4) Fe₂P-type (stable >~700 GPa) SiO₉

Tsuchiya & Tsuchiya (2011)

Ultrahigh-*P*,*T* phase relations in SiO₂



Tsuchiya & Tsuchiya (2011) PNAS

Experimental confirmation using a low-P analog, TiO₂



Dekura, Tsuchiya+ (2011) PRL

Experimental confirmation

- Laser shock technique
- Magnetic shock (Z-machine)

USA, France, ...

A new project in Japan at Spring-8



XFELとパワーレーザーによる 新極限物質材料の探索

> K. Tanaka, N. Ozaki, O. Sakata, T. Tsuchiya, T. Sano, T. Sekine, K. Arakawa, ... (>20 people)

Osaka Univ., NIMS, Ehime Univ., Hiroshima Univ., Shimane Univ., ...

X-ray Free Electron Laser (XFEL)

In situ observation in TPa regime

Significant optical absorption of Fe₂P-typ TiO₂

Dekura, Tsuchiya+ (2011) PRL





Electronic density of states

T Tsuchiya, 9th ISPS, 25 June 2012 Tsuchiya & Tsuchiya (2011) PNAS

Ag2p

Mg3s

Si3s

Si3p

Si3d 02s O2p

40



B2-MgO 1.5 TPa

B2-CaO 0.4 TPa





Band gap vs pressure in SiO₂ phases



Metsue & Tsuchiya (2012) Phys Chem Min

 Fe_2P -type SiO₂ has much smaller gap than Pyr but still remains insulating even above 1 TPa.

P2₁-type H₂O ice electronic DoS



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Hermann+ 2012 PNAS J. Tsuchiya, unpub.

Ice also still remains insulating even at 2 TPa.

Metallization in solids only for compounds with heavy (d) metals

Thermodynamic properties of Fe₂P-type SiO₂



T Tsuchiya, 9th ISPS, 25 June 2012 Modeling of the silica-rich planetary mantle

Density $\frac{d\rho}{dr} = \frac{\rho(r)g(r)}{\phi(r)}$ Gravity $\frac{dg}{dr} = 4\pi G\rho(r) - \frac{2Gm(r)}{r^3}$ Mass $\frac{dm}{dr} = 4\pi r^2 \rho(r)$ Pressure $\frac{dP}{dr} = -\rho(r)g(r)$

$$\phi(\mathbf{r}) = B_S(r) / \rho(r)$$

The ho - r relationship for a super-Earth with 10M $_{\oplus}$ evaluated by **Valencia+ (2006)**



Thermal structure of SE

Adiabatic temperature gradient

$$\left(\frac{dT}{dP}\right)_{S} = \frac{\alpha(P)g(P)T}{C_{P}(P)}$$

Almost no difference by using MgSiO₃ parameters


Phase transition buoyancy parameter

T Tsuchiya, 9th ISPS, 25 June 2012 Christensen and Yuen (1985)



Negatively large P_h



Small or positive P_h



Layered convection

Whole convection

For the Pyr-Cot transition in a super-Earth with $10M_{\oplus}$:

$$\Gamma = -10 \text{ MPa K}^{-1}$$

$$\frac{\Delta \rho}{\rho} = 0.04$$

$$\rho = 8400 \text{ kg m}^{-3}$$

$$\alpha = 0.53 \times 10^{-5} \text{ K}^{-1}$$

$$g = 31 \text{ m s}^{-2}$$

$$h = 4500 \text{ km}$$

$$\Rightarrow P_h \sim -0.064$$

The transition might have not so large effect even with a negative boundary.

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Decomposition of MgSiO₃ post-perovskite

Tsuchiya & Tsuchiya (2011)



 MgSiO₃ (PPv) → MgO (B2) + SiO₂ (Cot or Fe₂P) (Umemoto+ 2006; Tsuchiya & Tsuchiya 2011)

An intermediate state with MgSi₂O₅ + MgO (Umemoto+ 2011)

Disproportionation reaction in NaCoF₃ NaCoFe₃ (PPv) \rightarrow Na₅Co₃F₁₁ + NaCo₃F₇ (Yusa+ 2012)

Further studies with careful structure search maybe needed

But no eccentric changes such as metallization seem likely.

Energy transportation in the Earth



Lattice thermal conductivity

(a) **Phonon** = Quantized lattice vibrations



(b) Phonon-phonon scattering → Thermal resistivity

(c) Anharmonicity \rightarrow Interaction of phonons



Lattice thermal conductivity (Higher order anharmonic lattice dynamics)

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Previous works on κ

Ab initio (non-equilibrium) molecular dynamics (MD), etc

Large simulation cell size



Limited to simple crystal structures like MgO

(de Koker 2009; Tang & Dong 2010; etc)

Our technique

DFPT approach



Small (primitive) unit cell size

High efficiency & low numerical error

Applicable to complex structures like MgSiO₃

κ of MgSiO₃-Pv



Dekura, Tsuchiya, Tsuchiya, PRL, under review

Rayleigh number of mantles



Melting temperature

Ab initio two-phase coexisting MD

(Alfe 2009; Usui & Tsuchiya 2010; etc)

- Equilibrate a supercell with **Sol-Liq interfaces** at several P,T conditions
- A method to avoid the kinetic effects (super-cooling and -heating) across melting and freezing



a) T > T_M Melt stable Usui & Tsuchiya (2010) J Earth Sci

b) $T < T_M$ Solid stable



P = 73 GPa, T = 5,600 K

P = 62 GPa, T = 5,300 K

Melting curve of SiO₂



 T_M quite comparable to the core conditions of some planets

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Electronic DoS of silicate melt

Subsolidus condition (1800 GPa, 10000 K) Fe₂P-type SiO₂ Supersolidus condition (1800 GPa, 12000 K) Liquid SiO₂



Semi-metallic

Metallic

Metallization (band gap closure) across melting (cf. Karki 2007 PRB) Liquid silicate maybe easily mix with H or H_2O . \rightarrow **Core erosion (cf. Wilson+ 2012 PRL)**



Wilson & Militzer (2012) PRL

Oxides soluble to liquid H at $T_{gass giant}$ but not at T_{SE}

Current views

Gas & ice planets



- Molten metallic rocky core
 - Liquid iron core Active interior Core erosion

lacksquare

igodol

Super-Earths



- Solid insulating thermally well conductive rocky mantle
 - Solid iron core Less active interior

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Lattice thermal conductivity (Higher order anharmonic lattice dynamics)

$$\kappa = \frac{1}{3} \sum_{s}^{3n} \int \mathbf{v}_{\mathbf{q},s}^2 c_{\mathbf{q},s} \tau_{\mathbf{q},s} d\mathbf{q}$$

$$(\vec{0},s)$$
 (qj)
 $(-qj')$



Anharmonic phonon

Phonon lifetime

$$\tau_{\mathbf{q},s} = \frac{1}{2\Gamma(\omega_{\mathbf{q},s})}$$

Bose-Einstein function

Phonon damping function

$$\Gamma_{\mathbf{q}j}(\omega) = \frac{\pi}{2} \sum_{\mathbf{q}',j',j''} V_3(-\mathbf{q}j,\mathbf{q}'j',\mathbf{q}-\mathbf{q}'j'') \sum [1+n_{\mathbf{q}'j'} + n_{\mathbf{q}-\mathbf{q}'j''}] \delta(\omega_{\mathbf{q}'j'} + \omega_{\mathbf{q}-\mathbf{q}'j''} - \omega) + 2[n_{\mathbf{q}-\mathbf{q}'j''} - n_{\mathbf{q}'j'}] \delta(\omega_{\mathbf{q}'j'} - \omega_{\mathbf{q}-\mathbf{q}'j''} - \omega) + 2[n_{\mathbf{q}-\mathbf{q}'j''} - n_{\mathbf{q}'j'}] \delta(\omega_{\mathbf{q}'j'} - \omega_{\mathbf{q}-\mathbf{q}'j''} - \omega) \}.$$

 $n_{\mathbf{q}j} = \frac{1}{e^{\hbar\omega_{\mathbf{q}j}/k_BT} - 1}$ *V*₃: anharmonic coupling coefficient

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Bohr-Sommerfeld quantization condition



$$l = m_e vr = n\hbar$$

n = 1,2,3, ... \hbar : Plank constant

l is discrete, not constant.

$$\int \int pdq = n\hbar$$
Quantization condition

SCF cycle vs total energy variation

Au fcc structure



B2-CaO electronic DoS

T Tsuchiya, 9th ISPS, 25 June 2012 Tsuchiya & Tsuchiya (2011)





B2-CaO Vibrational DoS



Compression behaviors





Density variation of super-Earth with assuming pyrolitic composition

Periodic boundary condition for a crystal

$$\phi(x+a) = \phi(x)$$

$$Ga = \pm 2n\pi \ (n = 0, 1, 2, 3, \cdots)$$

$$G = 0, \pm \frac{2\pi}{a}, \pm \frac{4\pi}{a}, \cdots, \pm \frac{2n\pi}{a}$$

3-dimensional

$$\mathbf{G}_{\alpha} = \pm \frac{2n\pi}{\Omega} (\mathbf{a}_{\beta} \times \mathbf{a}_{\gamma})$$

Reciprocal lattice vector

第一原理電子状態計算の限界

①バンドギャップ問題: バンドギャップを過小評価



FIG. 3. Comparison of self-consistently calculated LDA and EXX band gaps (in eV) of various semiconductors with experimental data from Refs. 73 and 89–91.

Stadele+ (1999) PRB

②弱い結合:例)LDA結合距離を過小評価(overbind)

		Volume (Å ³)	Error (%)	Bulk Modulus (GPa)	Error (%)	Sublimation Energy (eV)	Error (%)
勾配補正 – (GGA)	LDA	26.43	-18	25.3	+132	0.99	+71
	BP86	30.85	-4	13.5	+24	0.68	+17
	PW91	31.35	-2	13.5	+24	0.55	-4
	PBE	31.82	-1	12.8	+17	0.53	-8
	B-Loc.	39	+22	4	-63	0.24	-42
	Exp.	32.05 ^a		10.9 ^b		0.58 ^c	

TABLE I. Properties of Bernal-Fowler ice.

Hamann+ (1997) PRB

GGAにより大きく改善

③遷移金属酸化物の基底状態(強相関電子状態)

金属電子状態



Cococcioni & de Gironcoli (2005) PRB

実際の絶縁基底状態(E_g ~2eV)が再現されない!

Melting curve of MgSiO₃



Comparable to the SiO₂'s T_M and also Fe's T_M

Study on the ultrahigh-pressure phases of Earth and planetary materials

- Just started and now rapidly progressing
- Unexpected phases are continuously discovered. There would still be many other unrevealed structures.
- Off-Hugoniot laser shock appears quite important to confirm calculations experimentally, maybe a unique technique from the experimental side.

Large scale computation

GRC-SRFC parallel cluster systems

Pyrope



Xeon

Knorringite1,2



World fastest







Observation of exsoplanets

Doppler method

Gravitational interaction between a parent star and a planet

- \mathbf{V}
- Existence of a planet
 - Mass of a planet



Transit method

Eclipse by a planet, i.e., the light of a parent star dimmed by a transiting planet

 \downarrow

- Existence of a planet
 - Size of a planet

Mass + Size \Rightarrow Mean density



Sasselov (2008)

Eutectic melting relation in the MgO-SiO₂ system



Lattice Dynamics method

- Atomic thermal vibration in sold
- ⇒ Collective motion of oscillators (phonon)





Harmonic Approximation $\Delta E=(k/2)\Delta x^2$ Linear Approximation F=-k∆x

Dynamical matrix



$$\sum \Phi_{\kappa\kappa'}^{\alpha\beta}(0l) \exp\{-i\mathbf{q} \cdot (\mathbf{x}_0 - \mathbf{x}_l)\}$$

Phonon dispersion relation

A milestone from K (November/2011): First-principles calculations of electron states of a silicon nanowire

www.riken.jp





100,000 atoms!!!

Calculations with a few hundred atoms now not special

Examples of the transit planet

CoRoT-7b

GJ1214b



Discovery: 2009 $M = \sim 4.8 M_{\oplus}$ $R = \sim 1.7 R_{\oplus}$ $\langle \rho \rangle = \sim 5.6 \text{ g/cm}^3 \sim \langle \rho_{\oplus} \rangle$

Discovery: 2009 $M = \sim 6.55 M_{\oplus}$ $R = \sim 2.7 R_{\oplus}$ $\langle \rho \rangle = \sim 1.9 \text{ g/cm}^3 < \langle \rho_{\oplus} \rangle$





Electronic structure of solids



Thermal structure of SE

Adiabatic temperature gradient

$$\left(\frac{dT}{dP}\right)_{S} = \frac{\alpha(P)g(P)T}{C_{P}(P)}$$




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