

Composition of Cometary Dust Particles Families of Comet 67P/ Churyumov-Gerasimenko

Martin Hilchenbach¹, Oliver Stenzel¹, John Paquette¹, Sihane Merouane¹, Henning Fischer¹, Klaus Hornung², Jouni Rynö³, Donia Baklouti⁴, Yves Langevin⁴, Hervé Cottin⁵, Nicolas Fray⁵, Robin Isnard⁵, Anais Bardyn⁶, Jochen Kissel¹

¹*Max-Planck-Institut für Sonnensystemforschung, Justus-von-Liebig-Weg 3, 37077 Göttingen, Germany,*
²*Universität der Bundeswehr LRT-7, Werner Heisenberg Weg 39, 85577 Neubiberg, Germany,*
³*Finnish Meteorological Institute, Erik Palmenin aukio 1, PO Box 503, 00101 Helsinki, Finland,*
⁴*Institut d'Astrophysique Spatiale, CNRS/Université Paris Sud, Bâtiment 121, 91405 Orsay, France,*
⁵*LISA, UMR CNRS 7583, Université Paris Est Créteil et Université Paris Diderot, Institut Pierre Simon Laplace, 94000 Créteil, France,*
⁶*Department of Terrestrial Magnetism, Carnegie Institution of Washington, 5241 Broad Branch Rd. NW, Washington DC, 20015, USA*

COSIMA, the COmetary Secondary Ion Mass Analyzer, is one of the three in-situ dust instruments onboard the Rosetta spacecraft, the ESA mission to comet 67P/Churyumov-Gerasimenko. From August 2014 to September 2016, Rosetta has been escorting the comet nucleus on its journey inwards and outwards the inner solar system. The instrument COSIMA collected cometary dust particles on metal targets in the inner coma, from 10 to hundreds of kilometers off the cometary nucleus. The targets are imaged and identified with an optical microscope and a selection of the collected particles are analyzed by secondary ion mass spectrometry (SIMS). Thousands of dust particles have been collected and the sizes of the identified particles and particle agglomerates range from 10 um up to sub-millimeter sizes. The mass spectra contain either positive or negative ions revealing both, mineral and organic, components. We will discuss the cometary dust particle composition and heterogeneity.