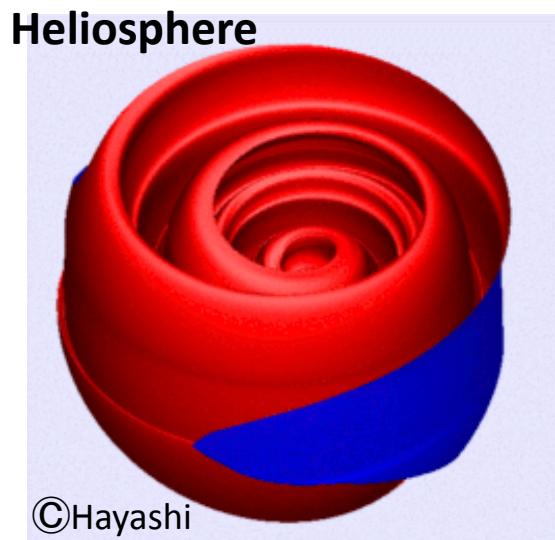


# 太陽圏システムと地球の気候/気象

宮原 ひろ子  
Hiroko MIYAHARA,

Institute for Cosmic Ray Research,  
The University of Tokyo, Japan

[hmiya@icrr.u-tokyo.ac.jp](mailto:hmiya@icrr.u-tokyo.ac.jp)



<http://sun.stanford.edu/~keiji/gallery11.html>

## Collaborators

Yusuke Yokoyama (AORI, The Univ. of Tokyo)  
Yasuhiko T. Yamaguchi (AORI, The Univ. of Tokyo)  
Takeshi Nakatsuka (Nagoya University)  
Hong K. Peng (The Univ. of Tokyo)  
Yukihiro Takahashi (Hokkaido University)  
Mitsuteru Sato (Hokkaido University)  
Hiroyuki Matsuzaki (MALT, The Univ. of Tokyo)  
Fuyuki Tokanai (Yamagata Univ)  
Yosuke Yamashiki (Kyoto Univ)  
Shuhei Masuda (JAMSTEC)  
John P. Matthews (Kyushu Univ)  
Kazuki Munakata (Shinshu Univ)  
Ryuho Kataoka (Tokyo Tech)

# 太陽活動と気候をとりまく諸問題

- ・太陽の放射強制力

太陽活動の長期変動にともなう総放射・スペクトル放射の絶対値、変動幅(daily～millennial)  
(気候モデルのvalidationへの影響)

- ・太陽活動が気候変動に影響するメカニズム

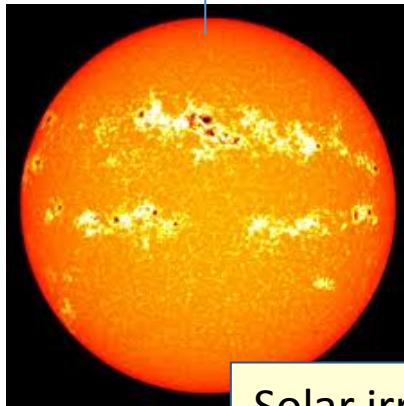
日射、スペクトル放射、太陽風、**宇宙線**(太陽圏磁場による減衰)

宇宙現象、地磁気強度変動と  
古気候との相関からも重要性が示唆されている

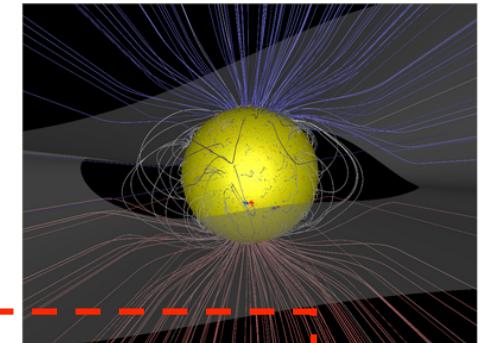
- ・太陽ダイナモ物理

11年変動、長期変動(88年周期、208年周期、1000年周期)  
27日周期の黒点・太陽フレア予測  
(最終的には、活動予測→気候予測モデルへのインプット)

27日



# Possible pathway of solar influence on climate change



## Solar activity

Solar irradiance  
Total / UV  
 $\sim 0.1\%$

Solar wind  
Particles  
(proton, etc.)

Solar/heliospheric  
magnetic field

$\sim 20\%$

Attenuation of cosmic rays

Promotion of  
chemical reaction  
in the stratosphere  
Ozone formation

Promotion of  
chemical reaction  
in the mesosphere  
NO<sub>x</sub> formation  
Ozone destruction

Nucleation

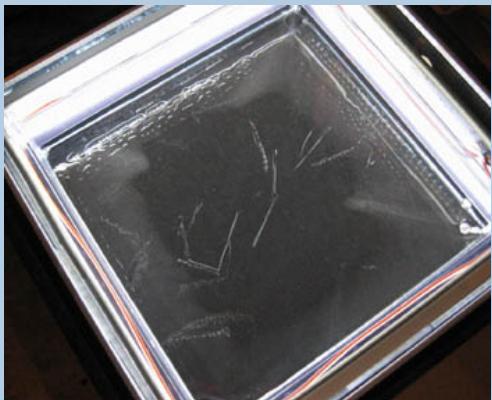
Polarize  
aerosols

Cloud property

Climate change

# Possible mechanisms

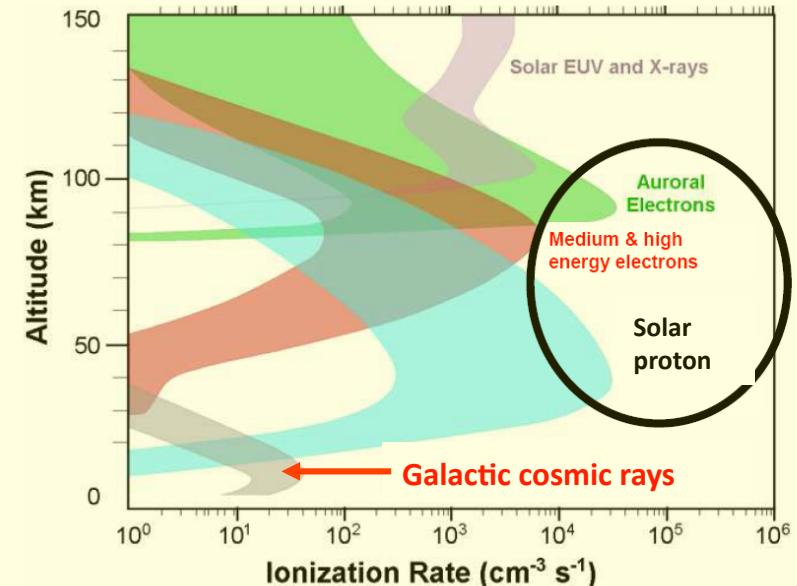
## 1. Nucleation



Formation of  
Cloud Condensation Nuclei

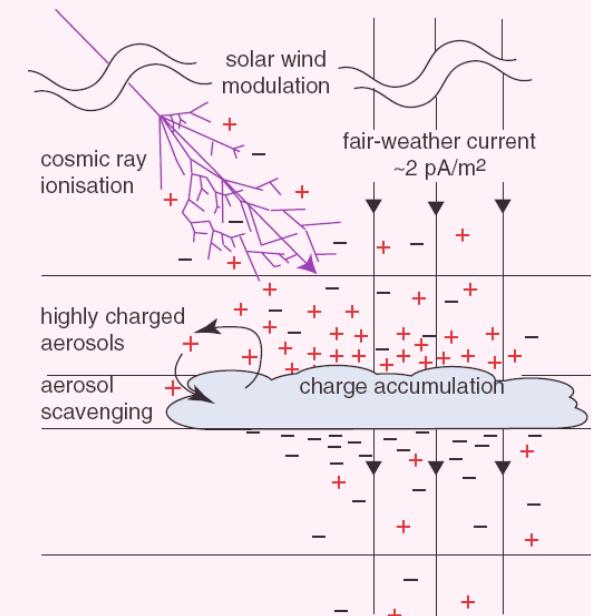
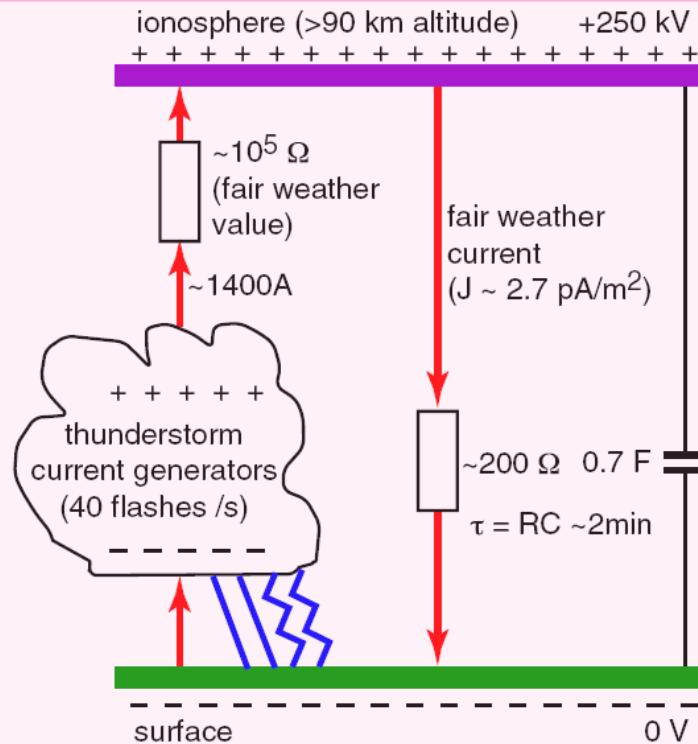
## 2. NO<sub>x</sub> formation

## Energetic Particle Precipitation

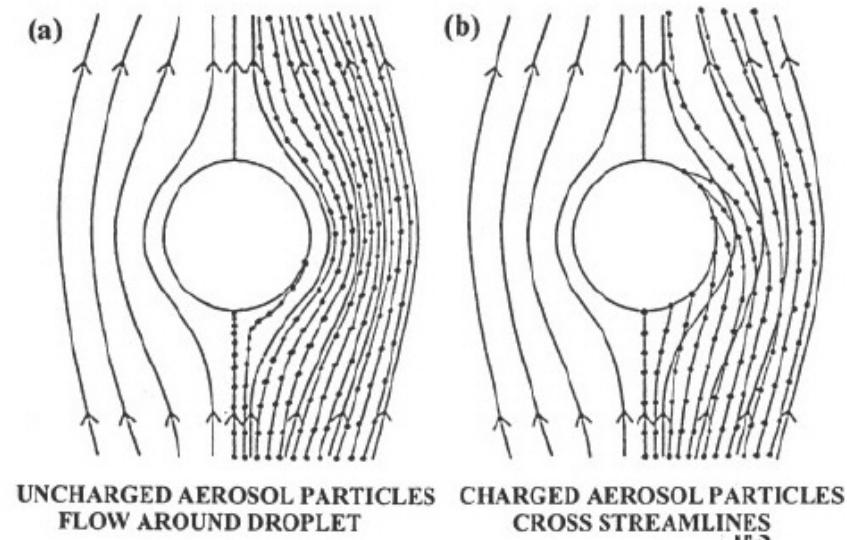


## 3. Coagulation

collision efficiency  
of aerosol particles



# 詳細な物理&定量的な議論は 今後の課題

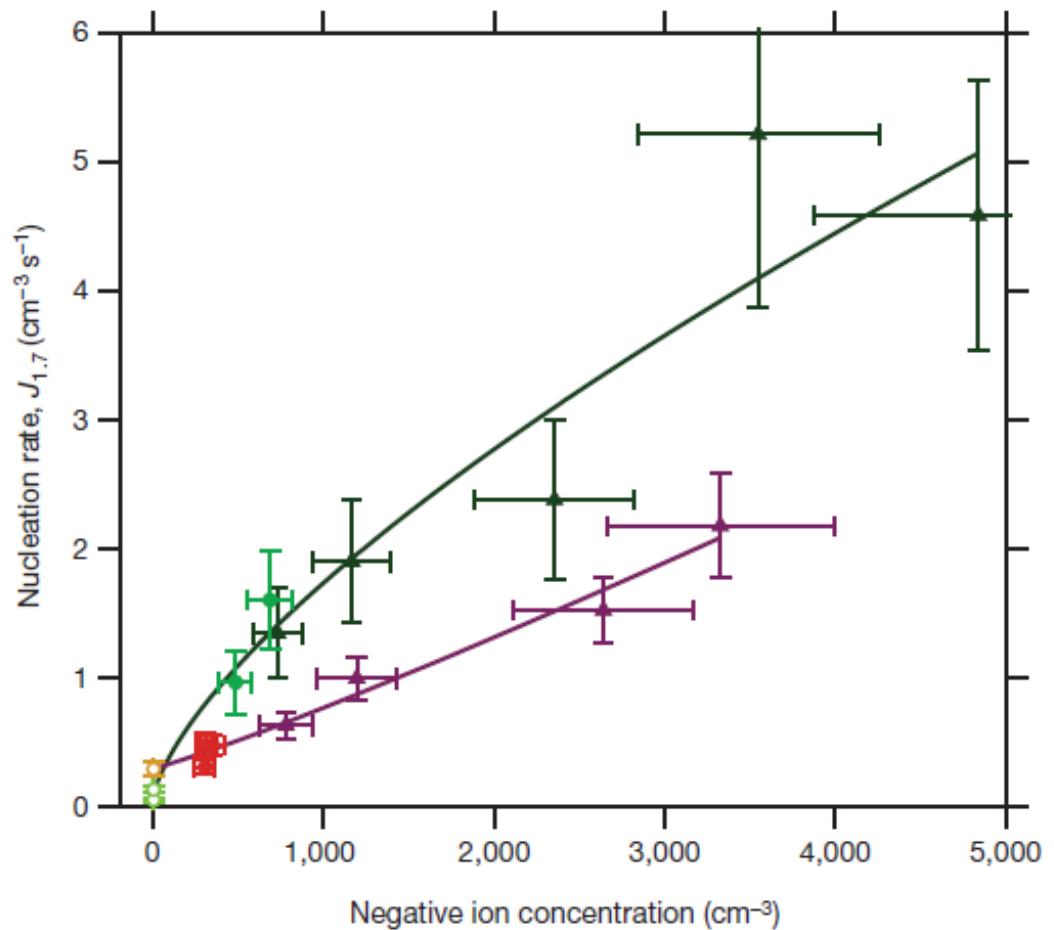


**Figure 5.3.** (a) Schematic of aerosol flow around a falling droplet in the absence of electrical forces. (b) Schematic of effect of electrical forces in moving aerosol particles across streamlines.

Tinsley 1996

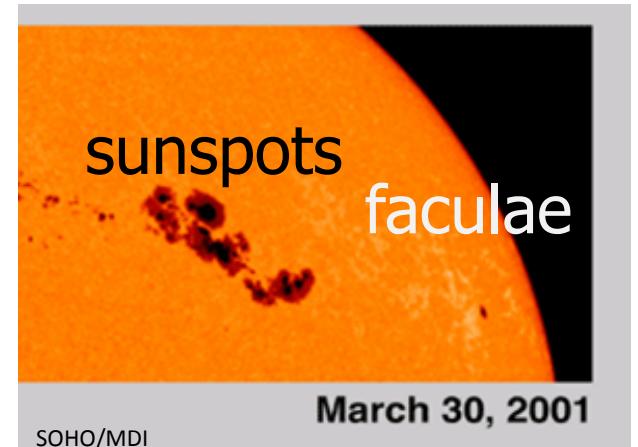
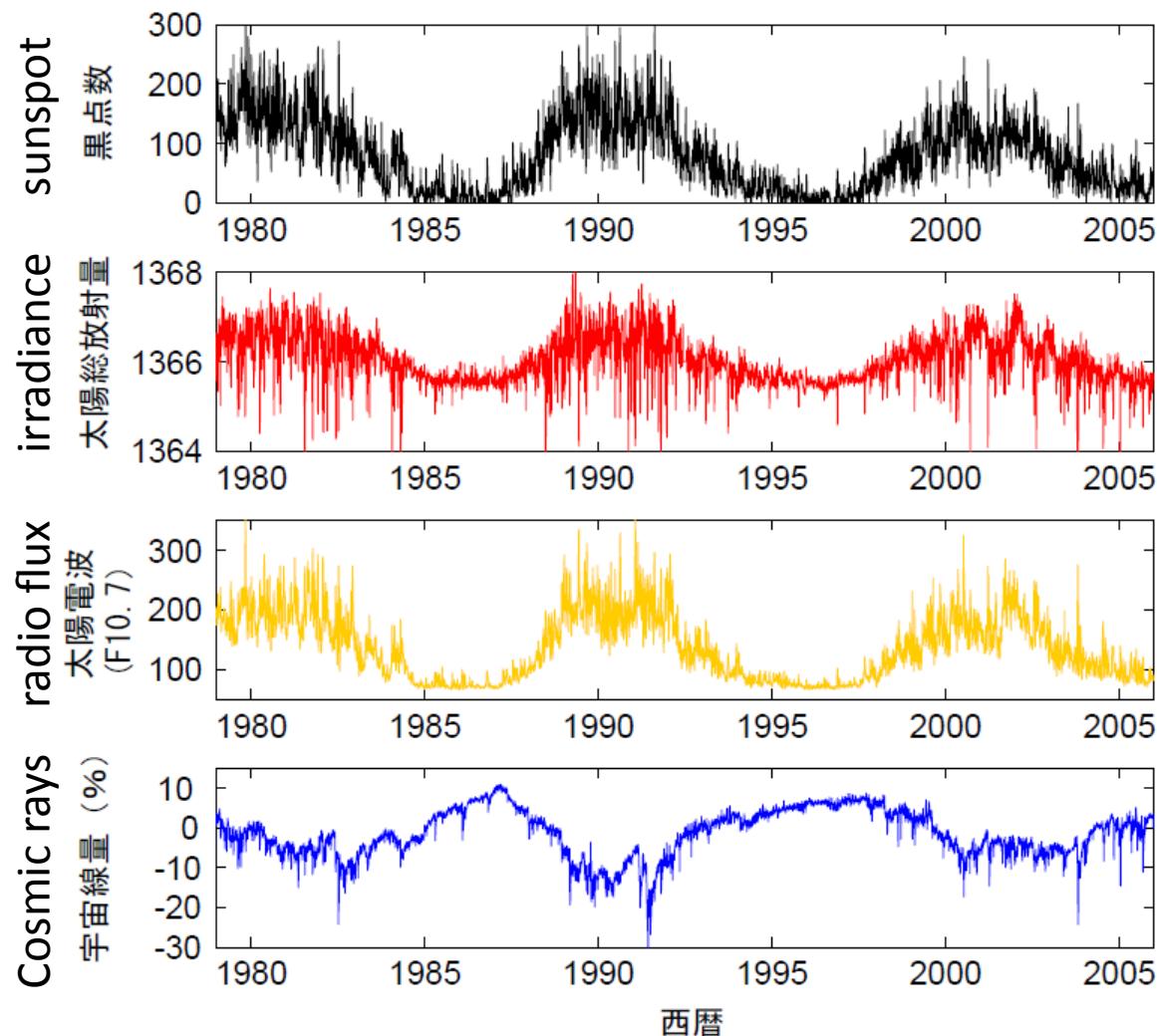
Tinsley & Yu 2006

Kirkby, Nature, 2011



**Figure 2 | Plots of nucleation rate against negative ion concentration.** Nucleation rates as a function of negative ion concentration at 292 K and  $[\text{H}_2\text{SO}_4] = 4.5 \times 10^8 \text{ cm}^{-3}$  (purple line), and at 278 K and  $[\text{H}_2\text{SO}_4] = 1.5 \times 10^8 \text{ cm}^{-3}$  (green line). Triangles,  $J_{\text{ch}}$ ; filled circles,  $J_{\text{gcr}}$ ; open circles,  $J_{\text{n}}$ . All measurements were made at 38% relative humidity and 35 p.p.t.v.  $\text{NH}_3$ . Neutral nucleation rates,  $J_{\text{n}}$ , were effectively measured at zero ion pair concentration (ion or charged-cluster lifetime  $< 1 \text{ s}$ ). The curves are fits of the form  $J = j_0 + k[\text{ion}^-]^p$ , where  $j_0$ ,  $k$  and  $p$  are free parameters. The error bars indicate only the point-to-point  $1\sigma$  errors; the nucleation rates and ion concentrations each have estimated overall scale uncertainties of  $\pm 30\%$ .

# Eleven-year variation in solar related parameters



現代においては、いずれも同期しており影響の区別をつけにくい → より長期的な変動の理解、古気候との比較

# Contents

太陽圏の磁場構造と宇宙線の22年周期変動

マウンダー極小期(西暦1645-1715年)における  
宇宙線変動と気候変動

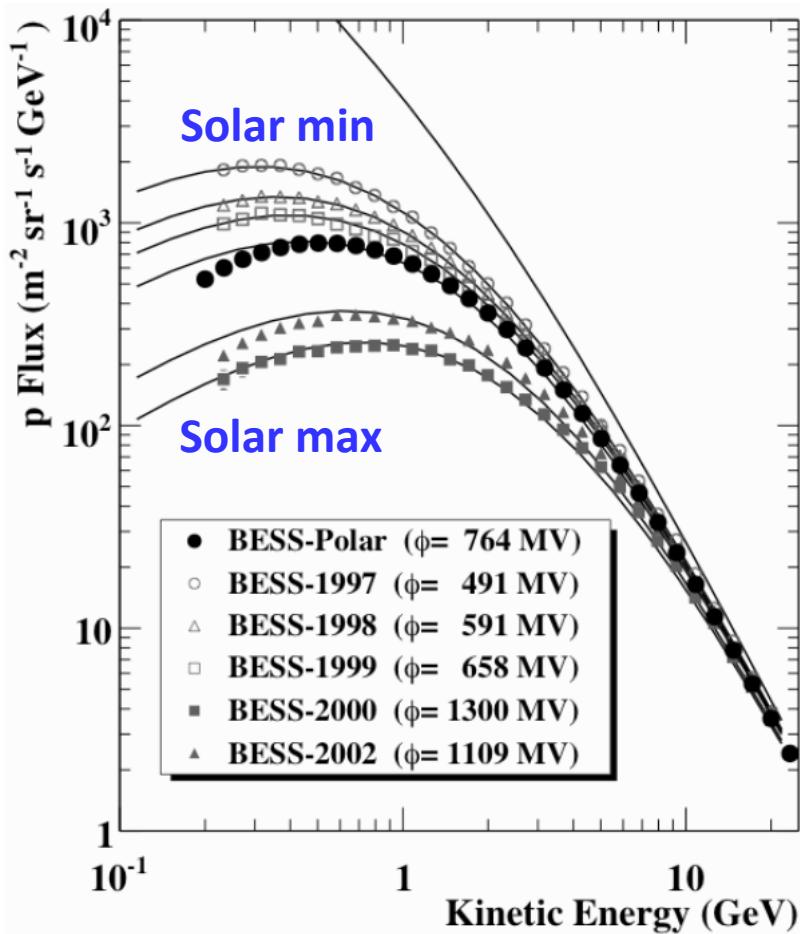
宇宙線の27日周期変動と赤道熱帯域の雲活動

(・宇宙線が気候システムに及ぼす変動のトレース

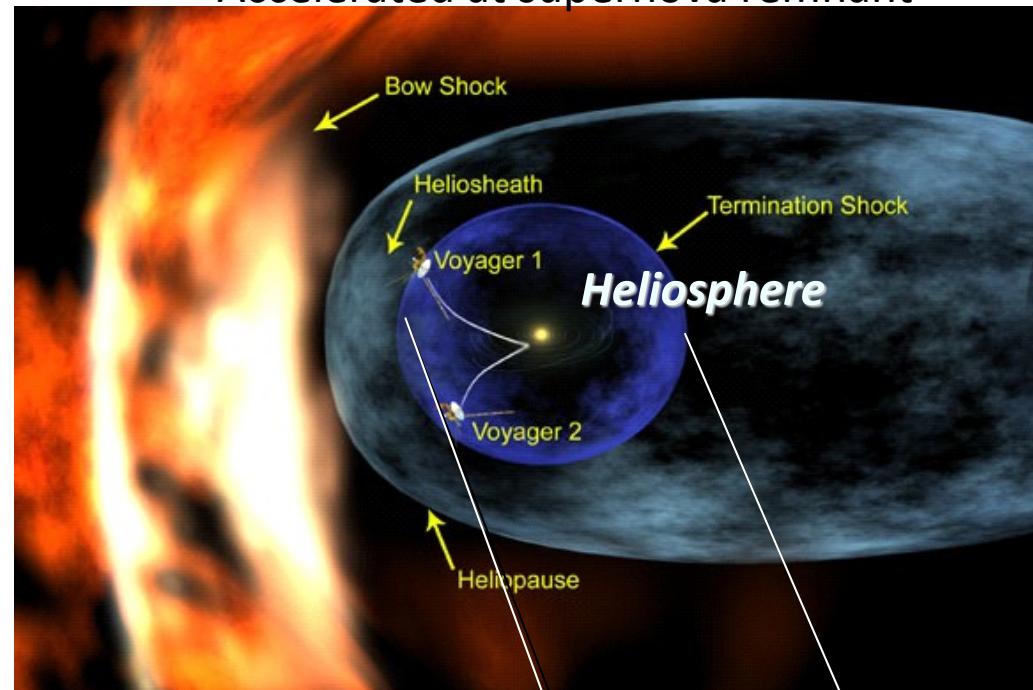
・気象への影響

の両観点から)

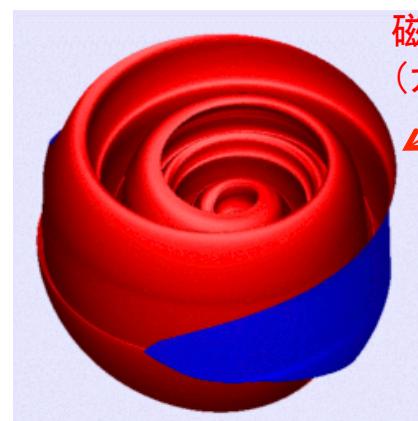
# Solar modulation of Galactic Cosmic Rays (GCRs)



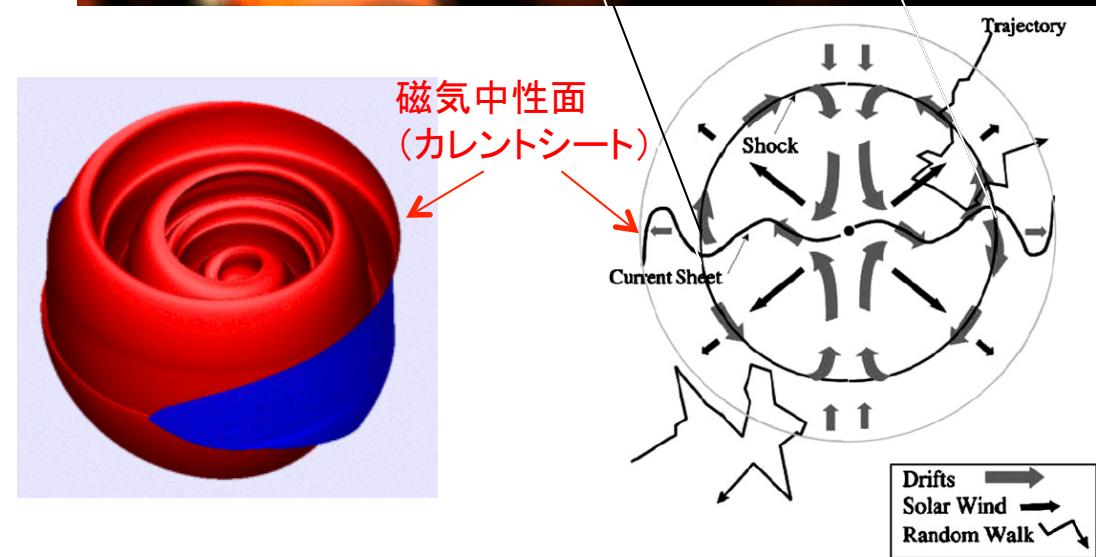
- Charged particles (mainly protons)
- Accelerated at supernova remnant



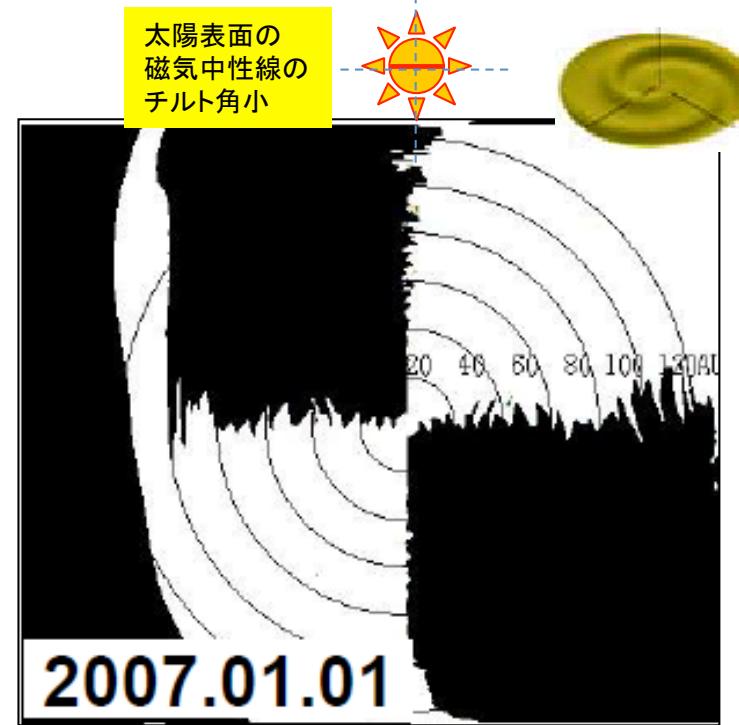
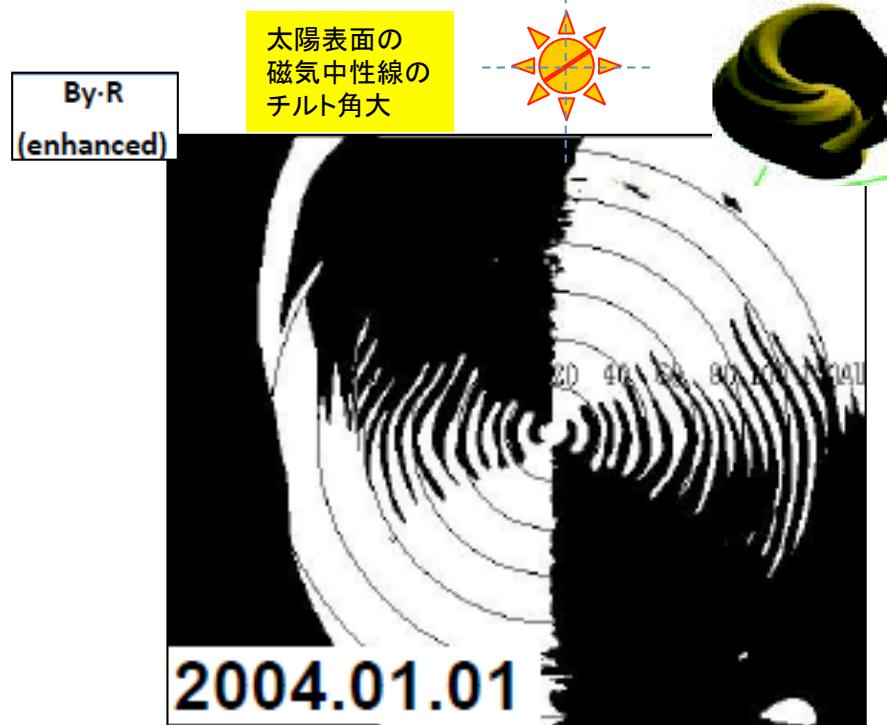
- diffusion
- advection by solar wind
- drift (  $B \times \nabla B$  ドリフト )



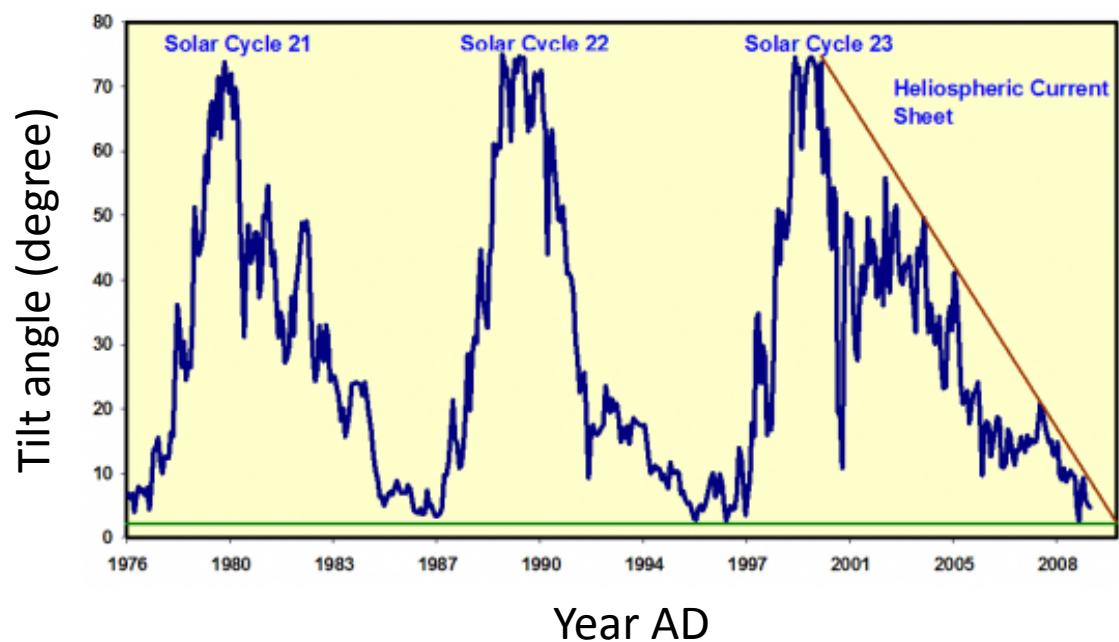
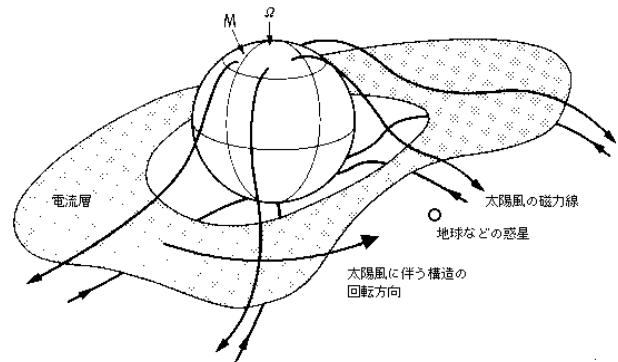
磁気中性面  
(カレントシート)



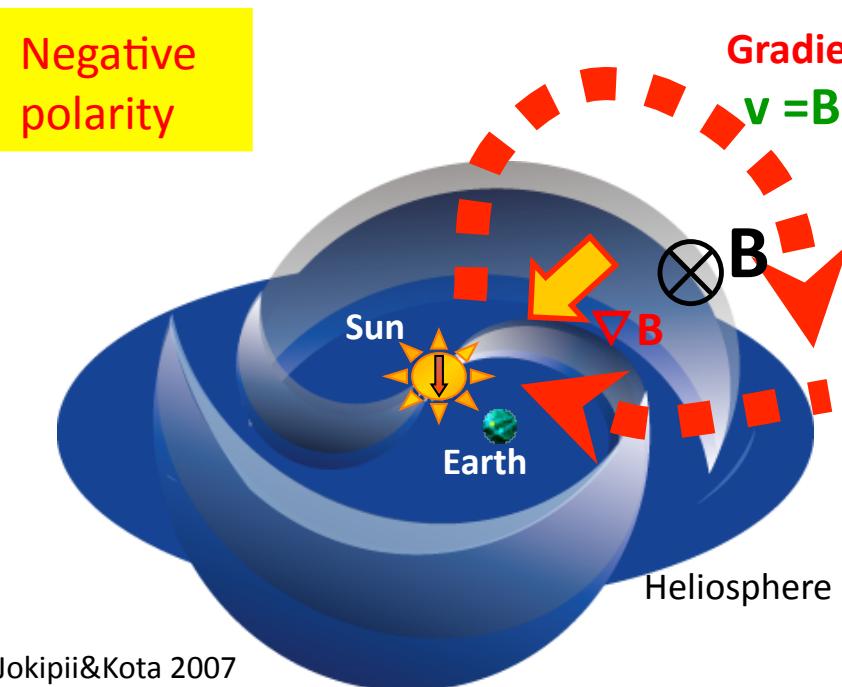
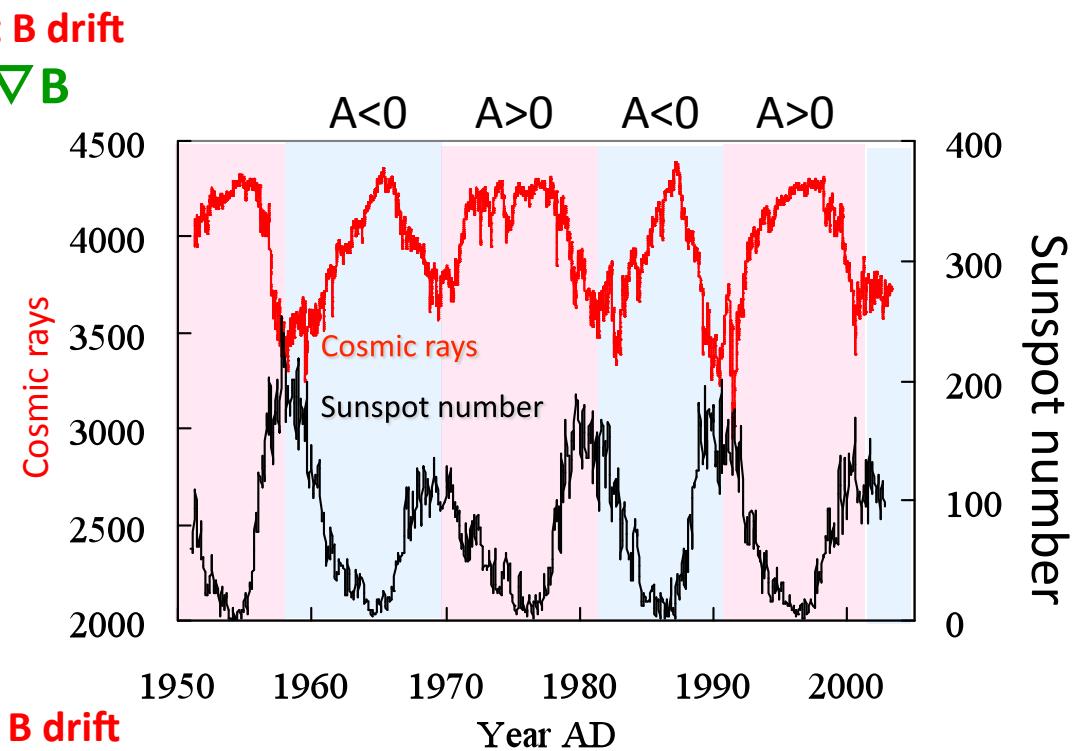
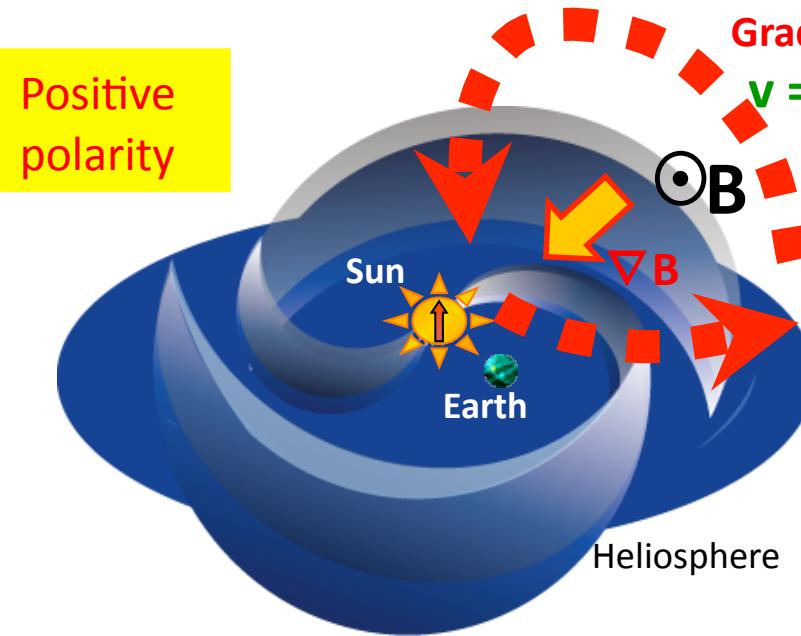
# Large scale structure of Heliospheric magnetic field



(based on Washimi et al)



# Cosmic ray variation & Solar magnetic polarity

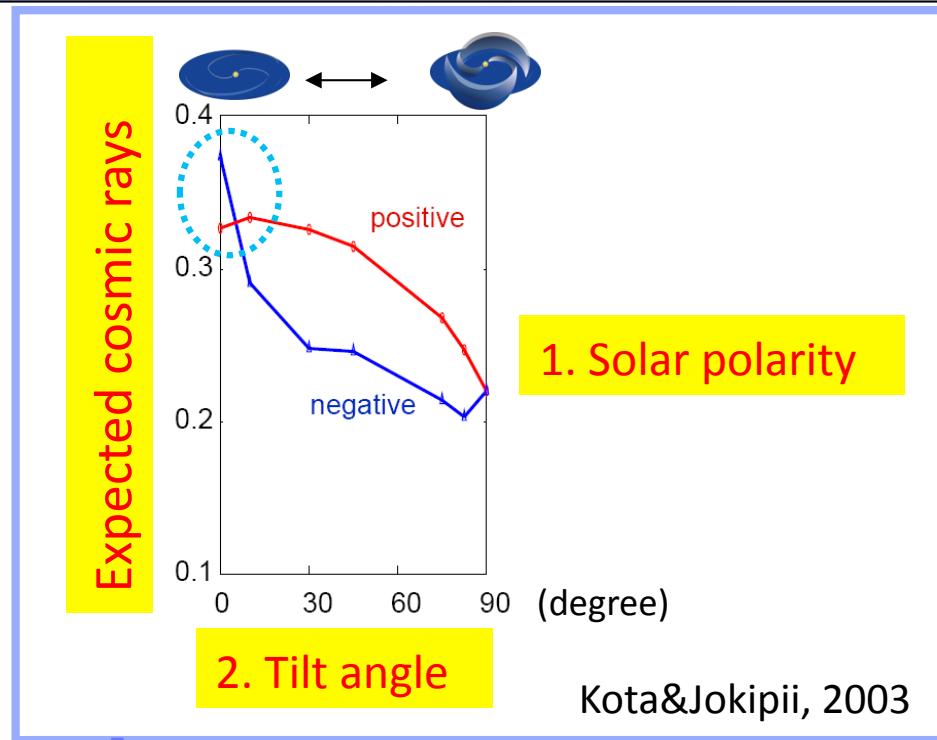


Important parameters for solar modulation

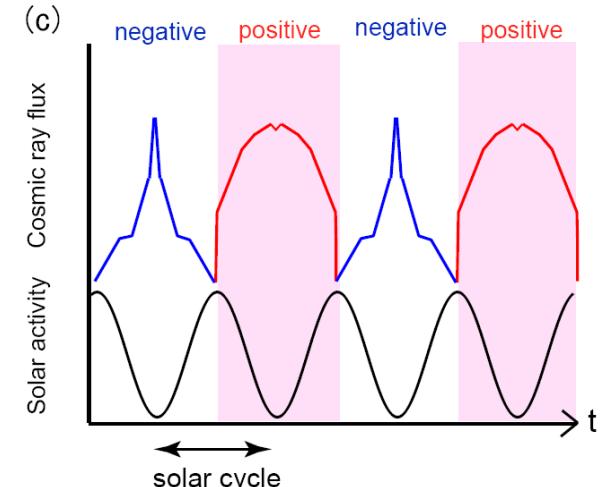
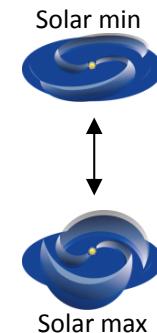
- solar dipole magnetic polarity
- tilt angle of heliospheric current sheet

# Variable “22-year” variation of cosmic rays

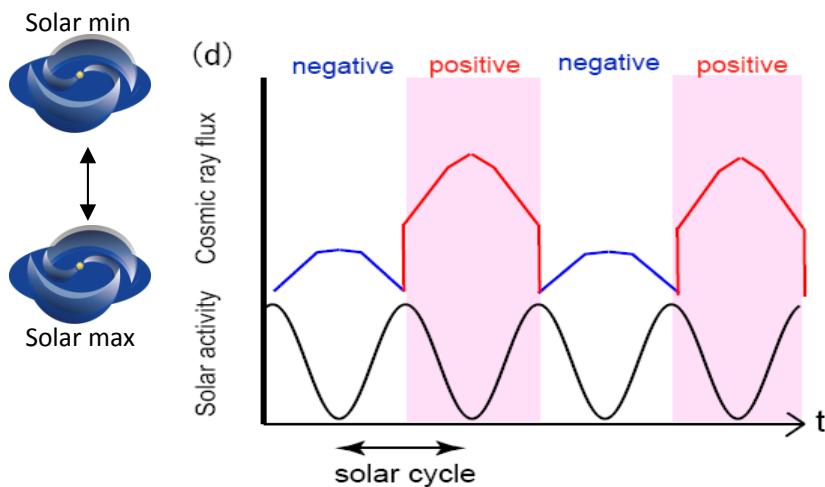
Miyahara et al., 2009



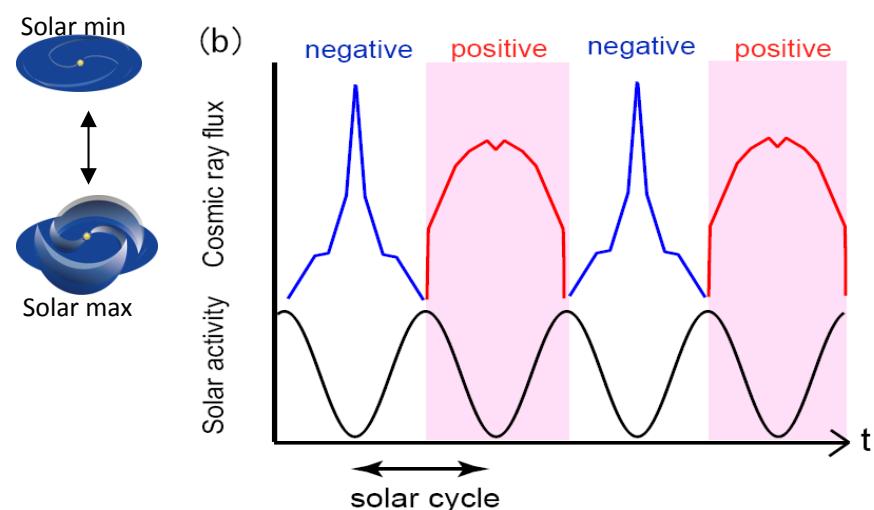
Modern: 5-75 degrees



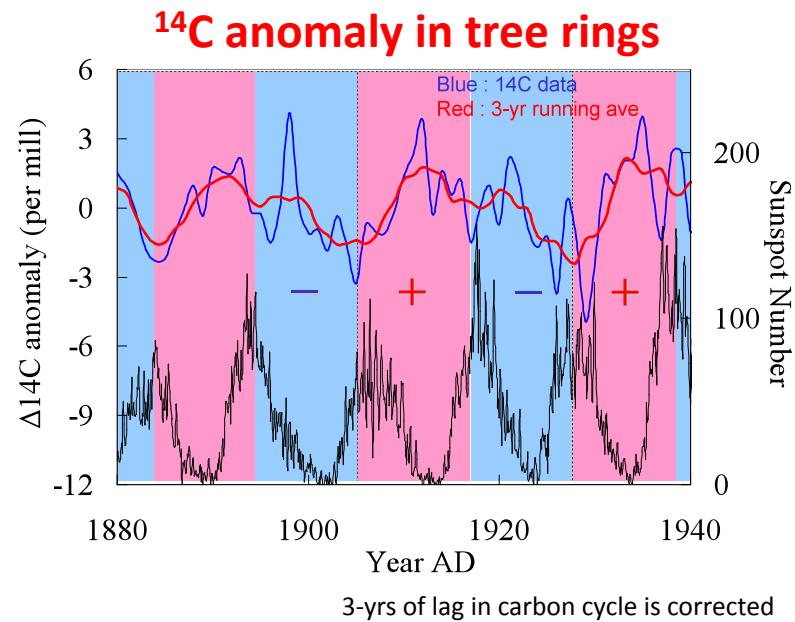
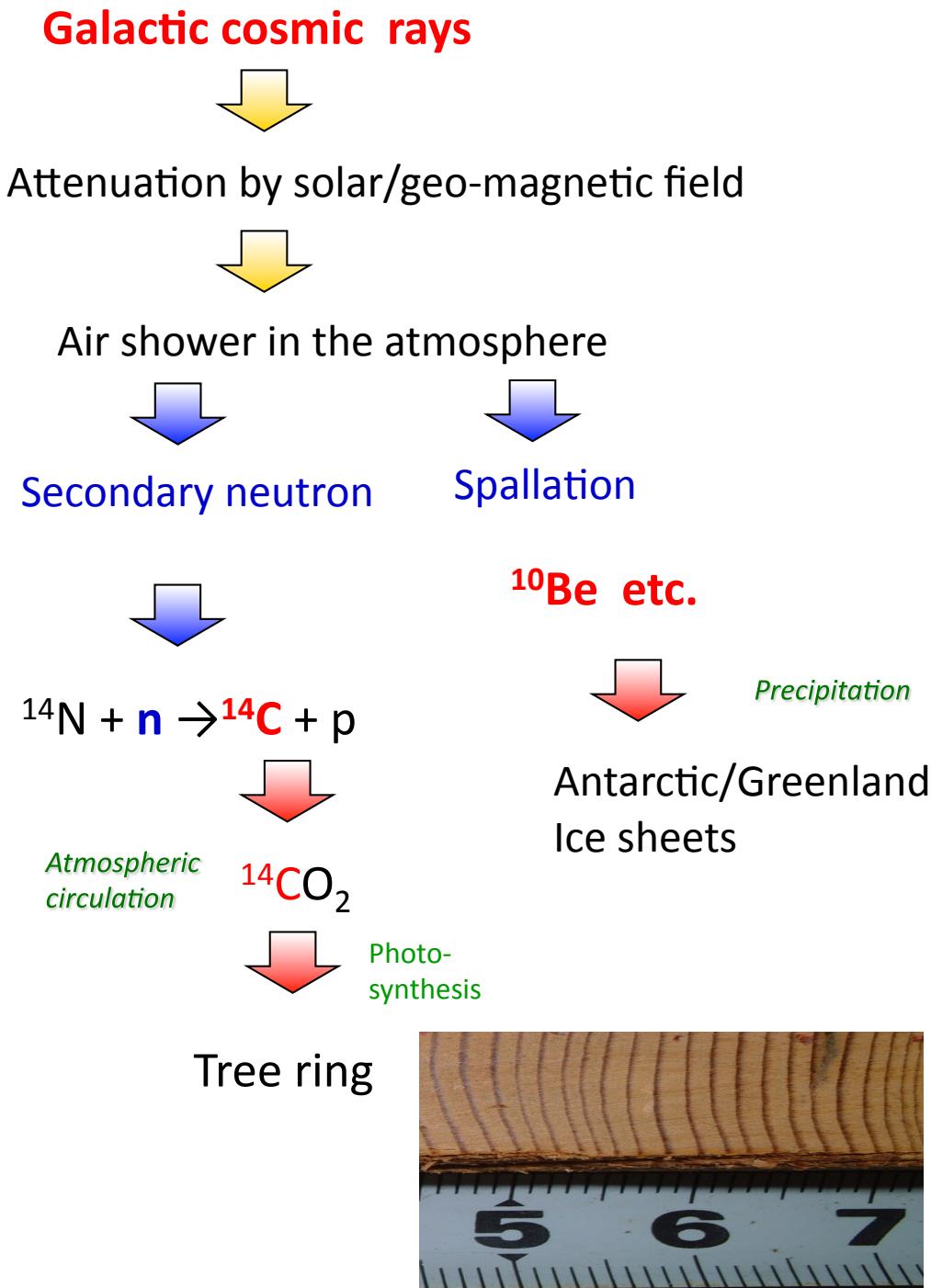
If 30-75 degrees



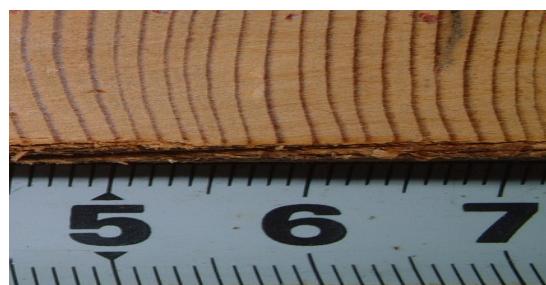
If 0-75 degrees



# Production of cosmogenic nuclides: $^{14}\text{C}$ and $^{10}\text{Be}$



Clear signal  
A few years of dating error



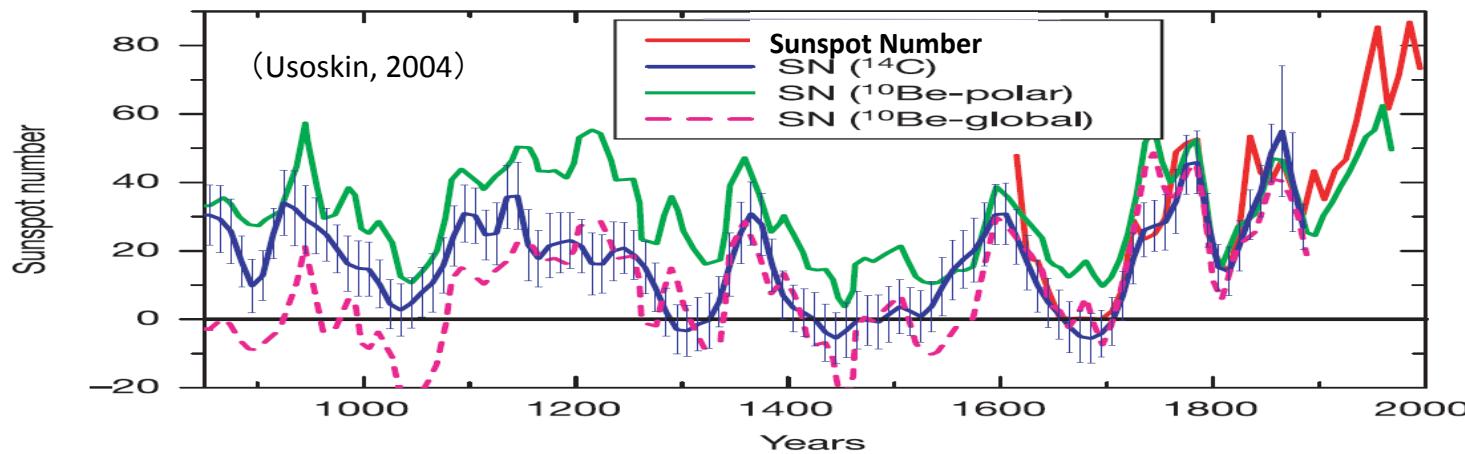
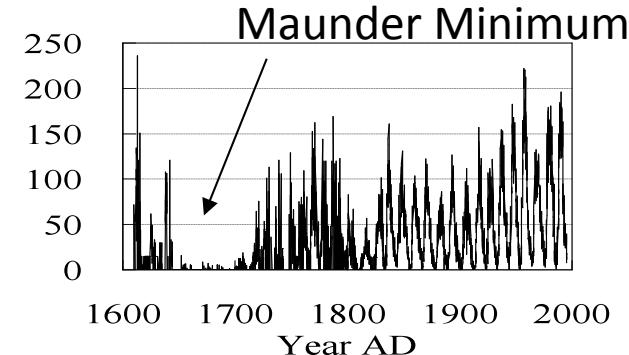
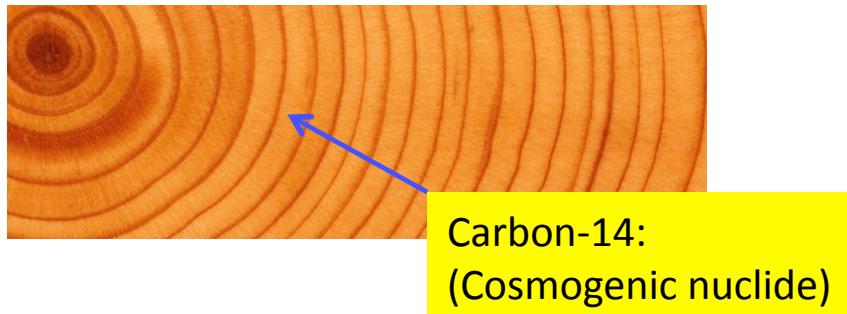
Absolute age  
Strongly attenuated signal

# Accelerator Mass Spectrometer

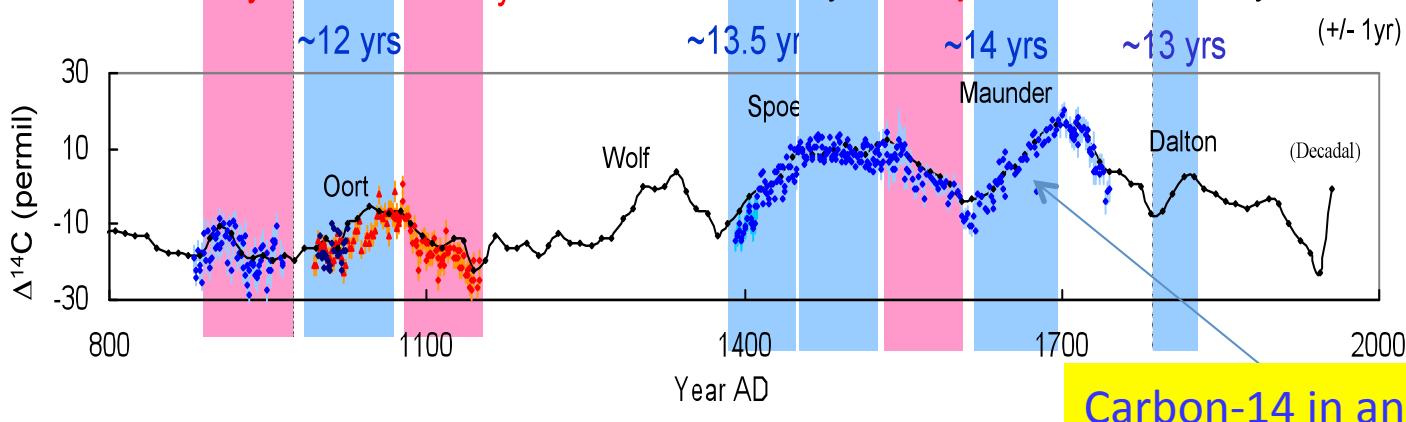
山形大学ホームページより



# Reconstructed solar decadal cycles in the past



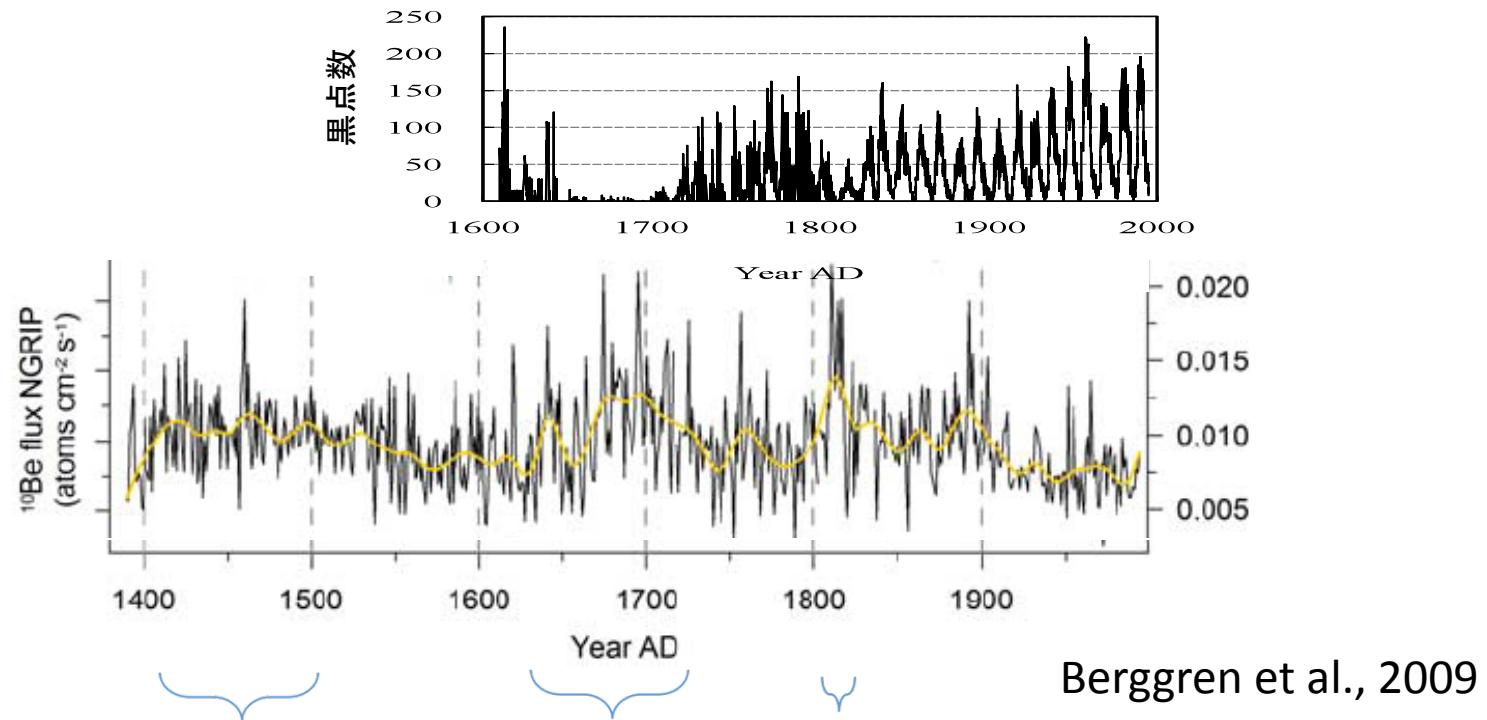
Actual mean length  
over the shaded  
period



Dacadal:  
Stuiver et al., 1998

Annual:  
Miyahara et al.,  
2004, 2006,  
2007, 2008

# グリーンランド氷床コア中のベリリウム10(宇宙線生成核種)から読み取る マウンダー極小期の宇宙線フラックス変動



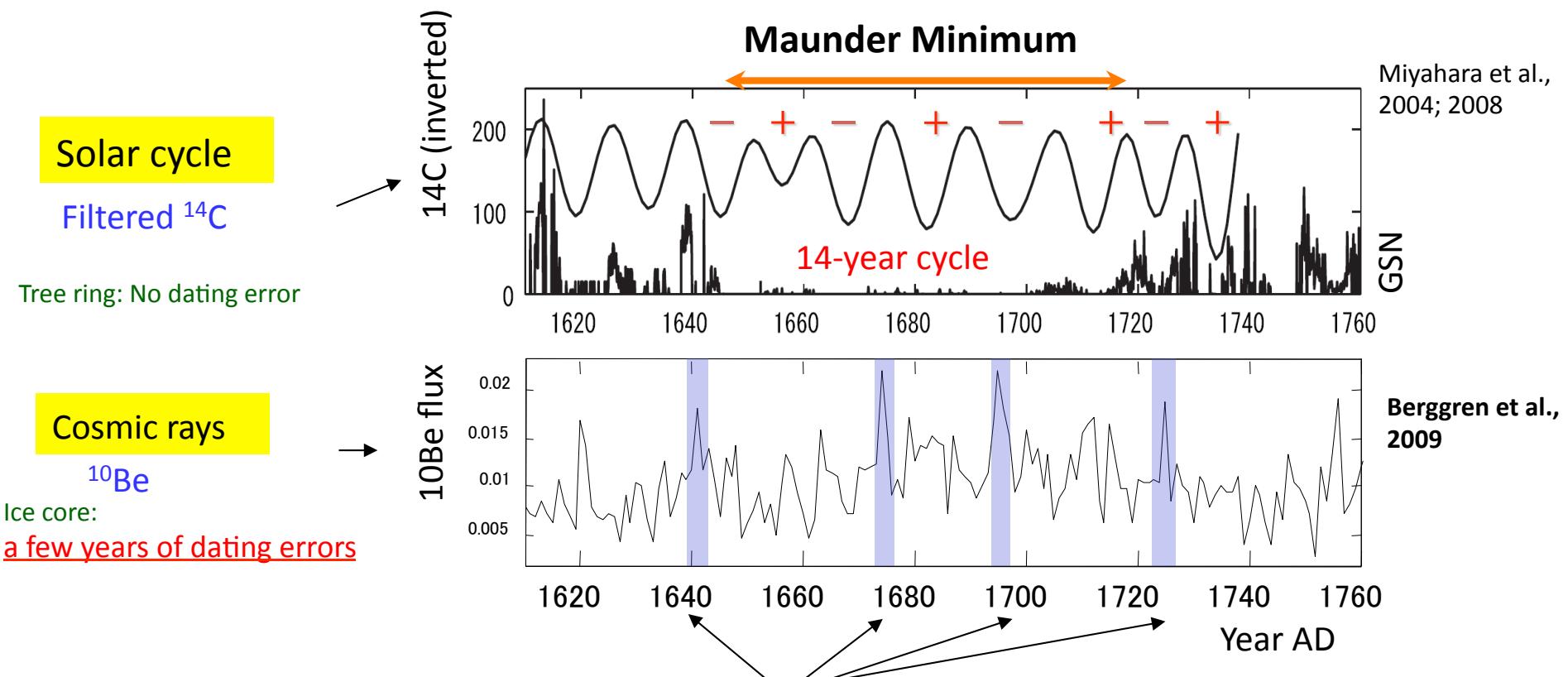
シュペーラー  
極小期  
(無黒点期)

マウンダー  
極小期

ダルトン  
極小期

# Cosmic-ray “22-year (28-year)” variation at the Maunder Minimum

Miyahara et al., IAU proc., 2009, Yamaguchi et al., PNAS, 2010



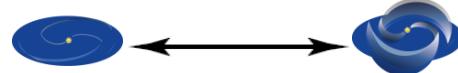
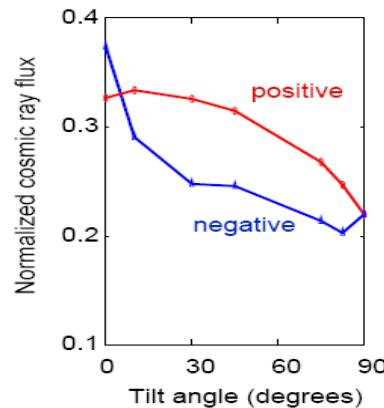
- Periodic cosmic ray enhancements, only for negative polarity ( $\sim 28$ -year period)
- **1-year scale enhancement, 30-50% higher than the peak for positive polarity**
- Significant manifestation of drift effect of cosmic rays in the heliosphere

# Pattern of cosmic ray variation at the Maunder Minimum and present

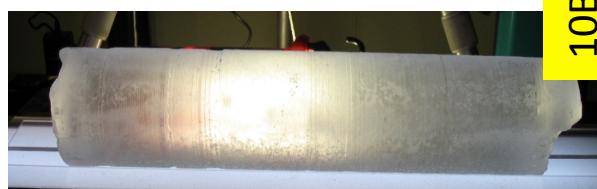
Miyahara et al., 2009

Based on  
Kota&Jokipii, 1983; 2003

- (a) 0 deg. at cycle min
- (b) 5 degs. at cycle min

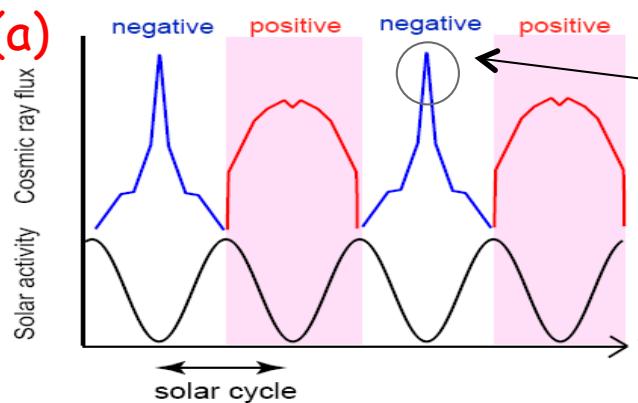


Heliospheric  
Magnetic field

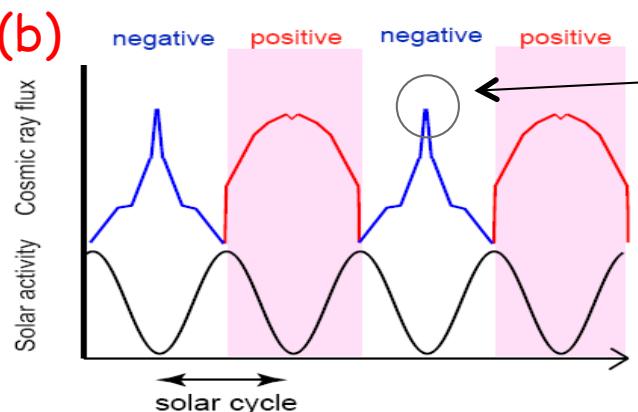


$^{10}\text{Be}$  flux

(a)

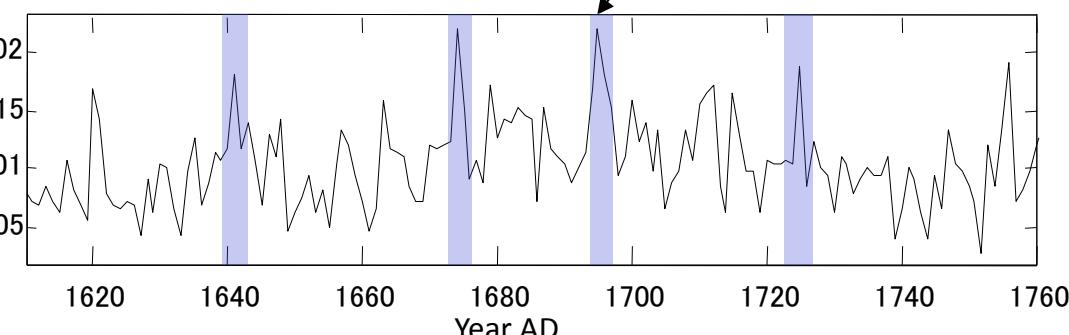


(b)



“Flattened current sheet model” reproduces  
 $^{10}\text{Be}$  variation

Berggren et al., 2009



# What $^{14}\text{C}$ and $^{10}\text{Be}$ suggests for the Maunder Minimum

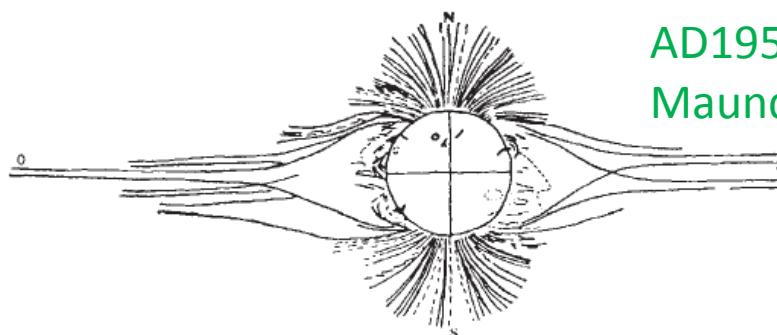
Solar Cycle length :  $\sim 14$  years

Magnetic polarity reversal : YES ( $\sim 28$ -year period)

Onset : two preceding 12-13 year cycles

Cosmic ray variations : Strong 22-year component

Heliospheric current sheet : More flattened

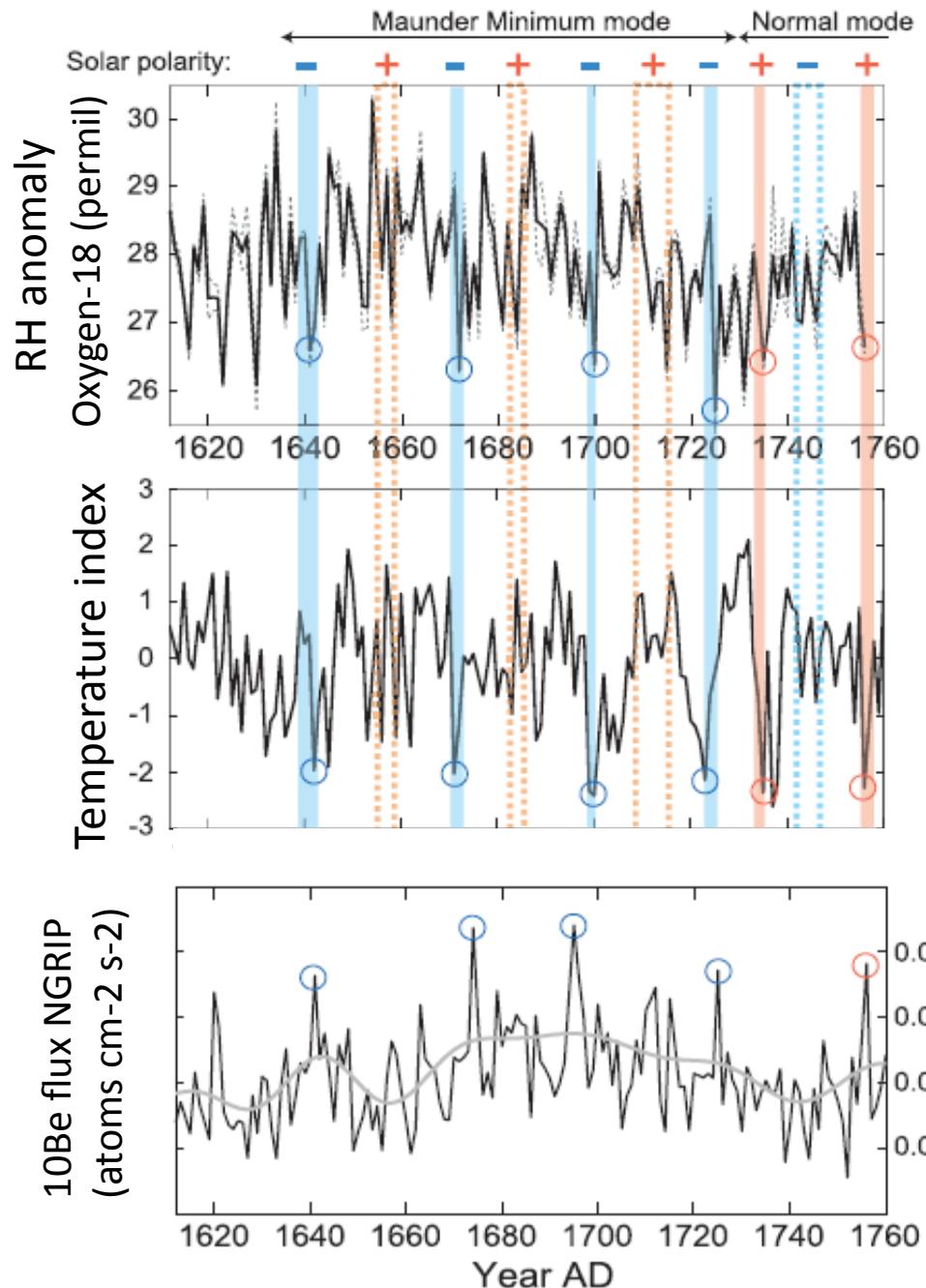


AD1954 case: stronger polar field  
Maunder Min: weaker equatorial field

Fig. 2 The structure of a sunspot minimum solar corona drawn from eclipse photographs<sup>11</sup> (June 30, 1954) obtained in Kozeletsk.

Any impact on climate?

# Climate response to cosmic-ray spikes during the Maunder Minimum



Yamaguchi, Yokoyama,  
Miyahara et al., PNAS, 2010

Dry  
↓  
Humid

Warm  
↓  
Cold

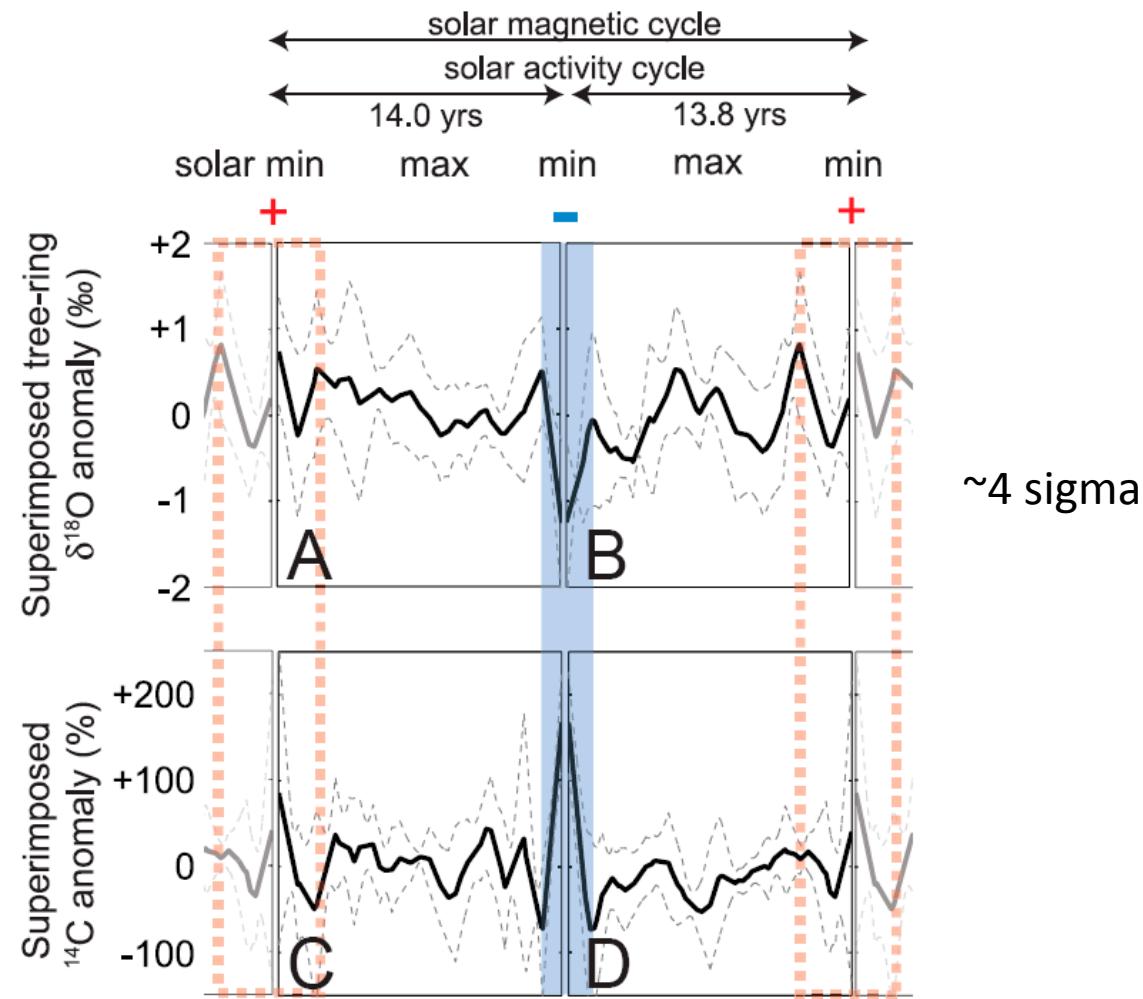
GCR flux  
(Berggren et al., 2009)

Ice core:  
a few years of dating errors

# Superposition of four 1-year spikes for $^{14}\text{C}$ (GCR) and $^{18}\text{O}$ (climate)

Relative Humidity  
anomaly  
reconstructed  
from tree rings

Tree ring  
 $^{14}\text{C}$  anomaly

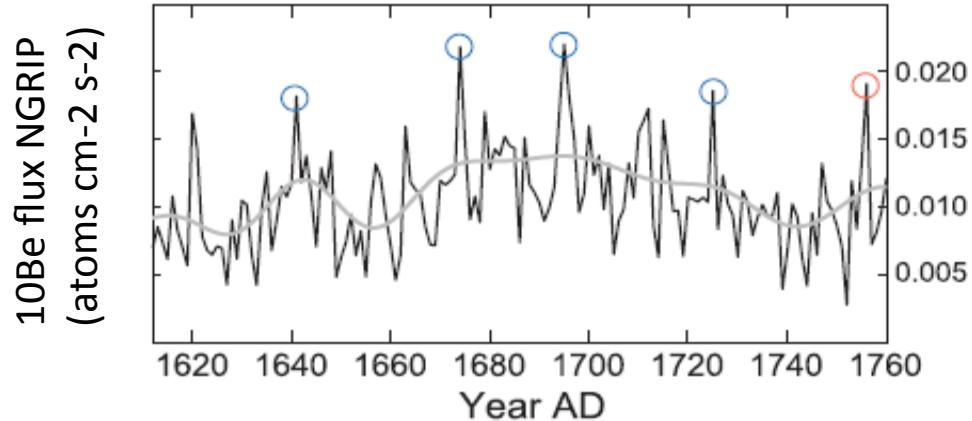


*No time lag!*

# 課題：宇宙線スパイクをトレーサーとした 全球気候応答マッピング

(1642年付近、1671年付近、1700年付近、1724年付近の計4イベント)  
→ 樹木年輪中炭素14濃度の超高精度分析で絶対年代決定

気温と降水の両方のマッピングにより、  
宇宙線に対する気候システムの複雑な応答の理解につなげたい



- ・降水の増減の地域性
- ・北半球(グリーンランド、日本)の  
気候変動の同期

## まとめ

- ・太陽圏システムとして地球気候、気象を捉えなおすことで気候の未解明の振動に関する理解が深まる可能性
  - ・太陽活動、磁場極性、太陽圏磁場構造、周期性などの変化について理解をより一層深め予測手法を確立する必要性
- 詳細な雲物理の解明については、雲の高度分布・粒径分布観測や室内実験に期待